



Trade-Off Between Fuel and Time Optimization

Shane Ross

Control and Dynamical Systems, Caltech

New Trends in Astrodynamics and Applications, Jan. 20-22, 2003

Interplanetary Mission Design

- *Use natural dynamics for fuel efficiency*
 - **Dynamical channels** connecting planets and moons
 - Trajectory generation using invariant manifolds in the 3-body problem suggests new numerical algorithms for interplanetary missions
- *How to balance fuel efficiency with reasonable flight times?*
 - Gravity assists, ballistic captures can take a long time
 - Short flight times important for challenging missions, e.g. *Multi-Moon Orbiter* to multiple Jovian moons

Multi-Moon Orbiter

- *Orbit each moon in a single mission*
- Other Jovian moons are also worthy of study
 - All may have oceans, evidence from *Galileo* suggests

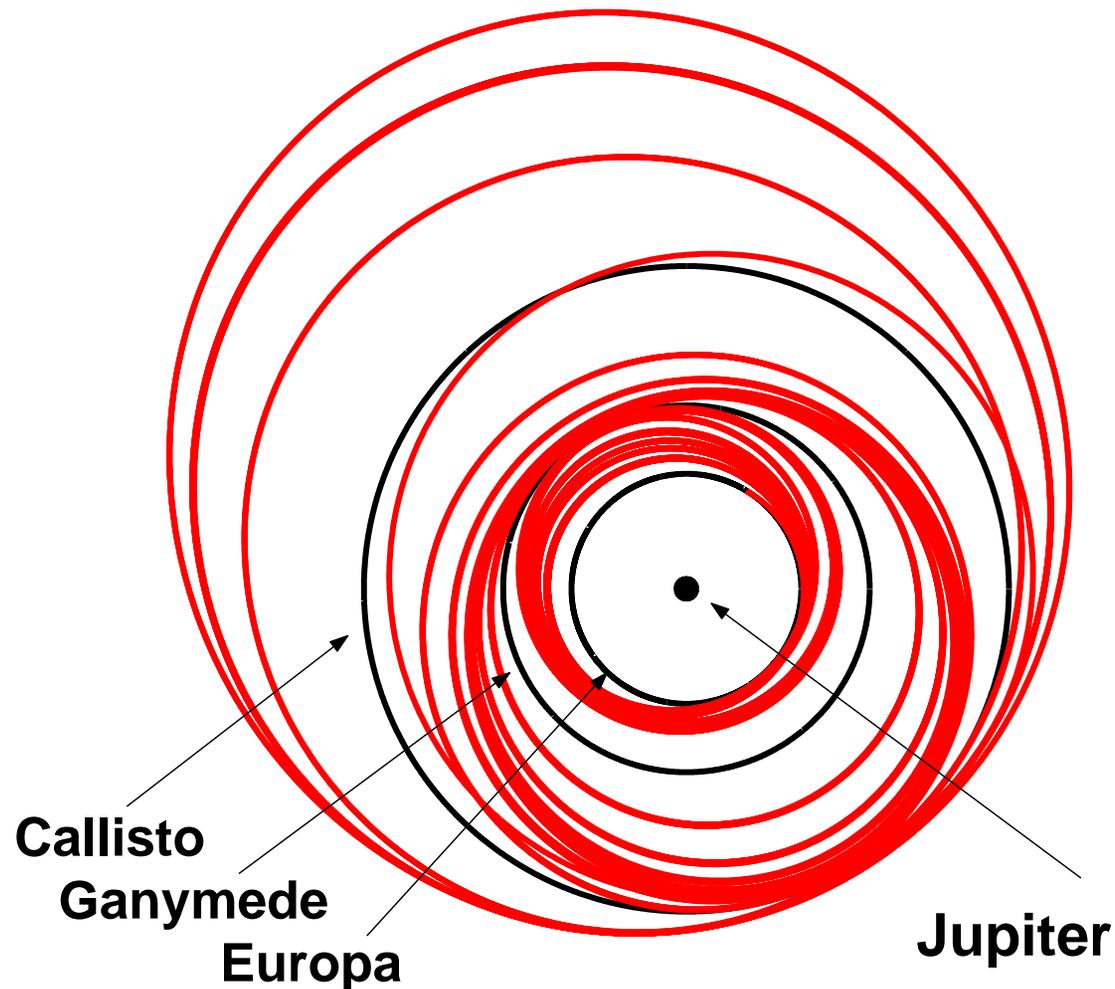


Multi-Moon Orbiter

- ΔV is low (~ 20 m/s), but flight time ~ 4 years

Low Energy Tour of Jupiter's Moons

Seen in Jovicentric Inertial Frame



Fuel vs. Time Trade-Off

■ *Motivating Example*

■ *Earth to Moon Trajectories*

- Consider a transfer from Earth orbit to lunar orbit
- Previously addressed by Belbruno and Miller, where nonlinear n -body effects were used to lower ΔV at the expense of a longer time of flight compared to Hohmann
- A good example problem for seeking efficient numerical algorithms to find fuel vs. time trade-off, applicable for many other situations

Fuel vs. Time Trade-Off

- Use planar circular restricted 3-body model
 - Consider the effect of only the Earth and Moon
- Compare with earlier methods; Hohmann, Bolit and Meiss, Schroer and Ott
 - Hohmann: simple and fast, but fuel-expensive
 - Remove recurrent loops in a chaotic trajectory
 - Target passes between mean motion resonances
- We will give some background on these methods

Some Background

■ *Starting/final orbital parameters*

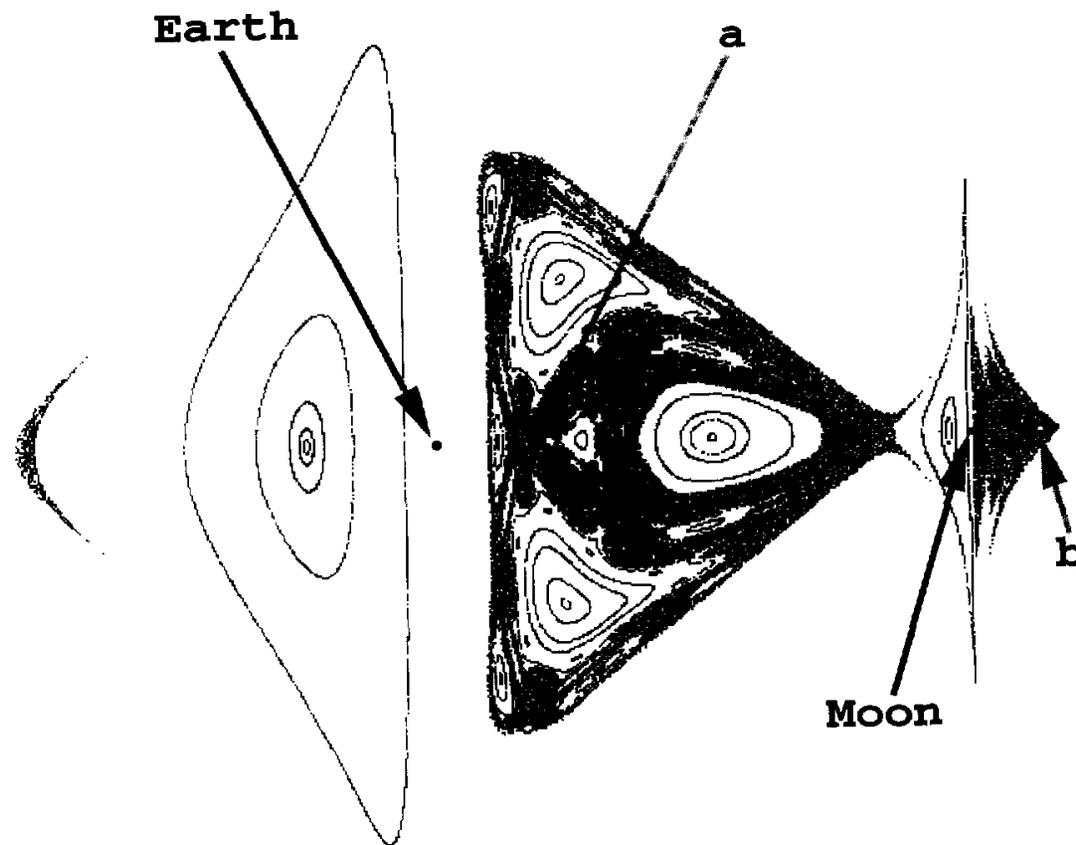
- Starting orbit: 59559 km circular Earth orbit
- Final orbit: lunar orbit with perilune 13970 km

■ *Transfer trajectory*

- Classical method: Hohmann transfer
 - Two maneuvers: ΔV_1 and ΔV_2 , both large
- Total $\Delta V = 1220$ m/s, TOF = 6.6 days
- **First Goal:** use the same orbital parameters, but lower the total ΔV
- **Second Goal:** keep time of flight reasonable

Some Background

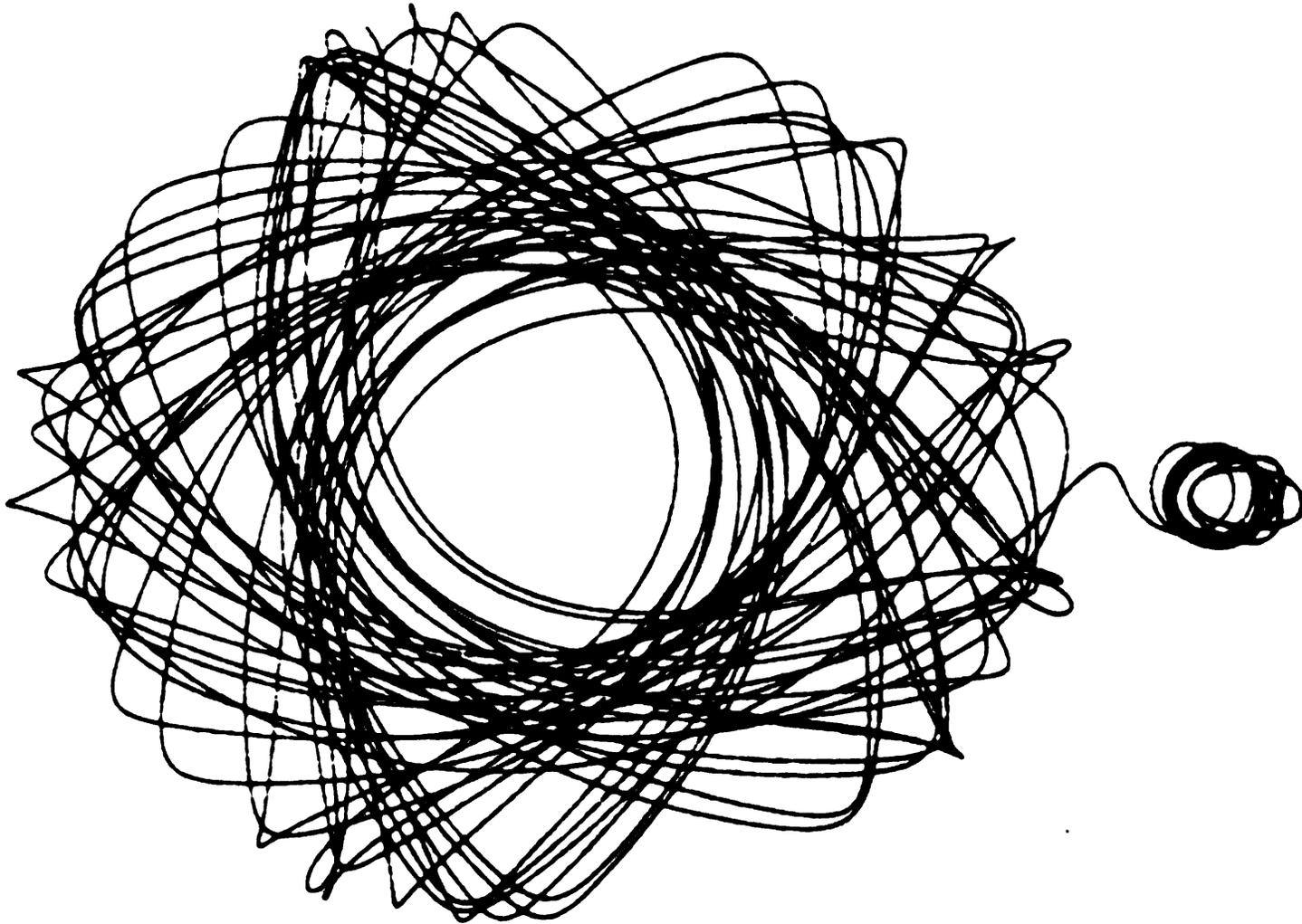
- Boltt & Meiss [1995] : target through recurrence
 - Find chaotic solution for fixed energy, remove recurrent loops with very small ΔV 's



Poincare section: chaotic and regular motion intermixed.

Some Background

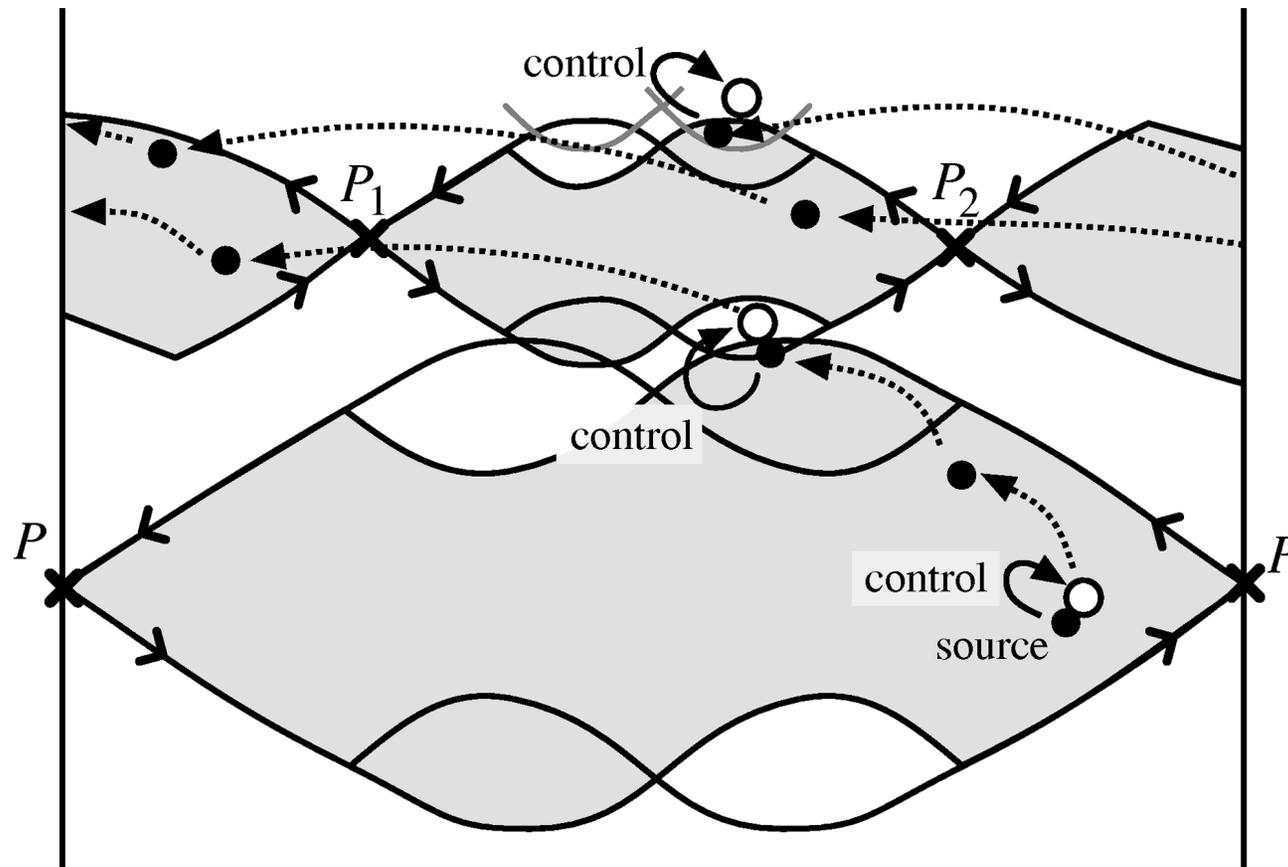
- Trajectory found : 750 m/s, 748 days



Spacecraft trajectory in the rotating frame

Some Background

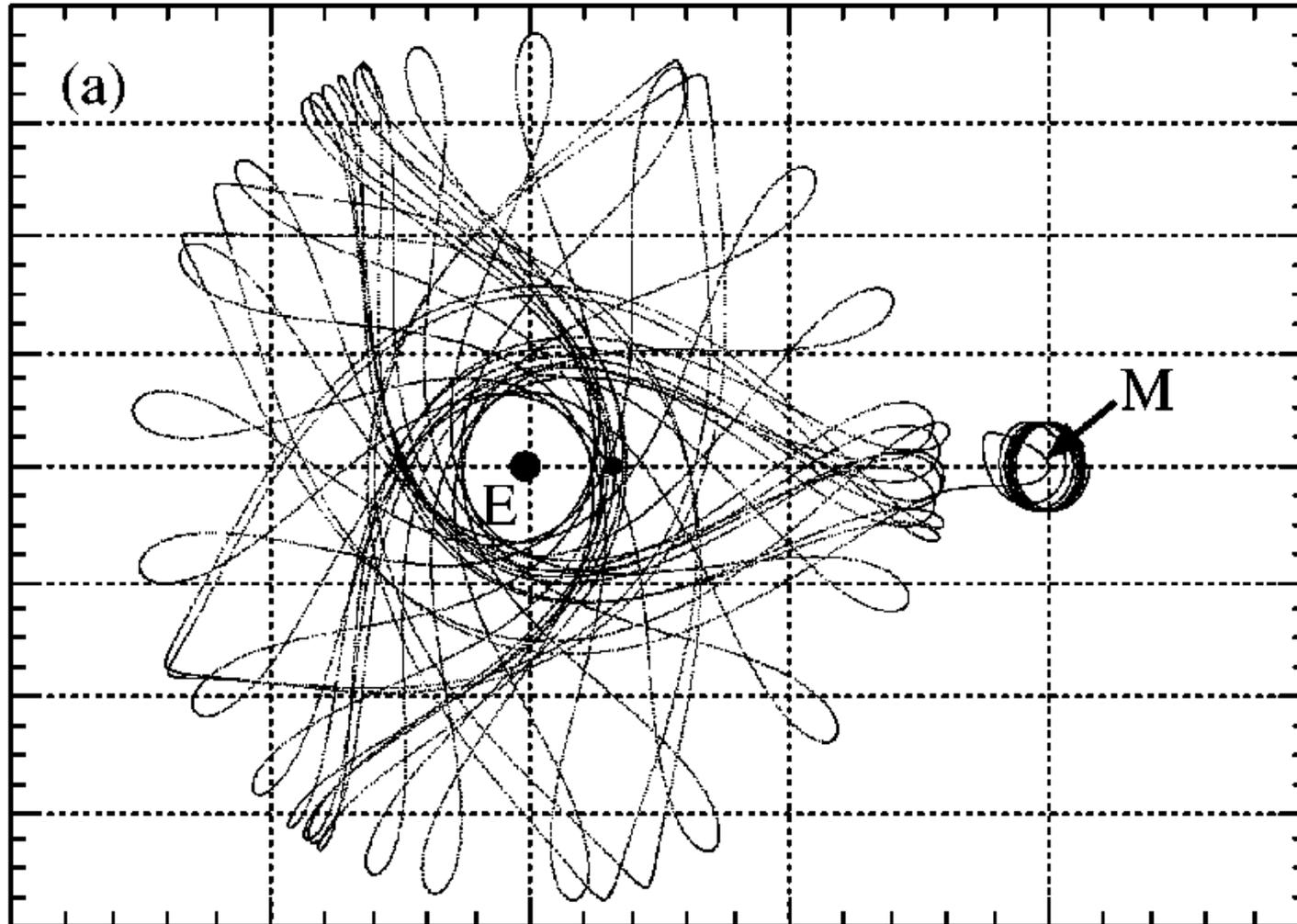
- Schroer & Ott [1997] : target passes btwn resonances
 - Leap-frog between a series of resonances using very small ΔV 's until trajectory reaches the moon



Resonances: target the passes between them with small controls.

Some Background

- Trajectory found : 749 m/s, 378 days



Spacecraft trajectory in the rotating frame

Our Approach

■ *Our approach*

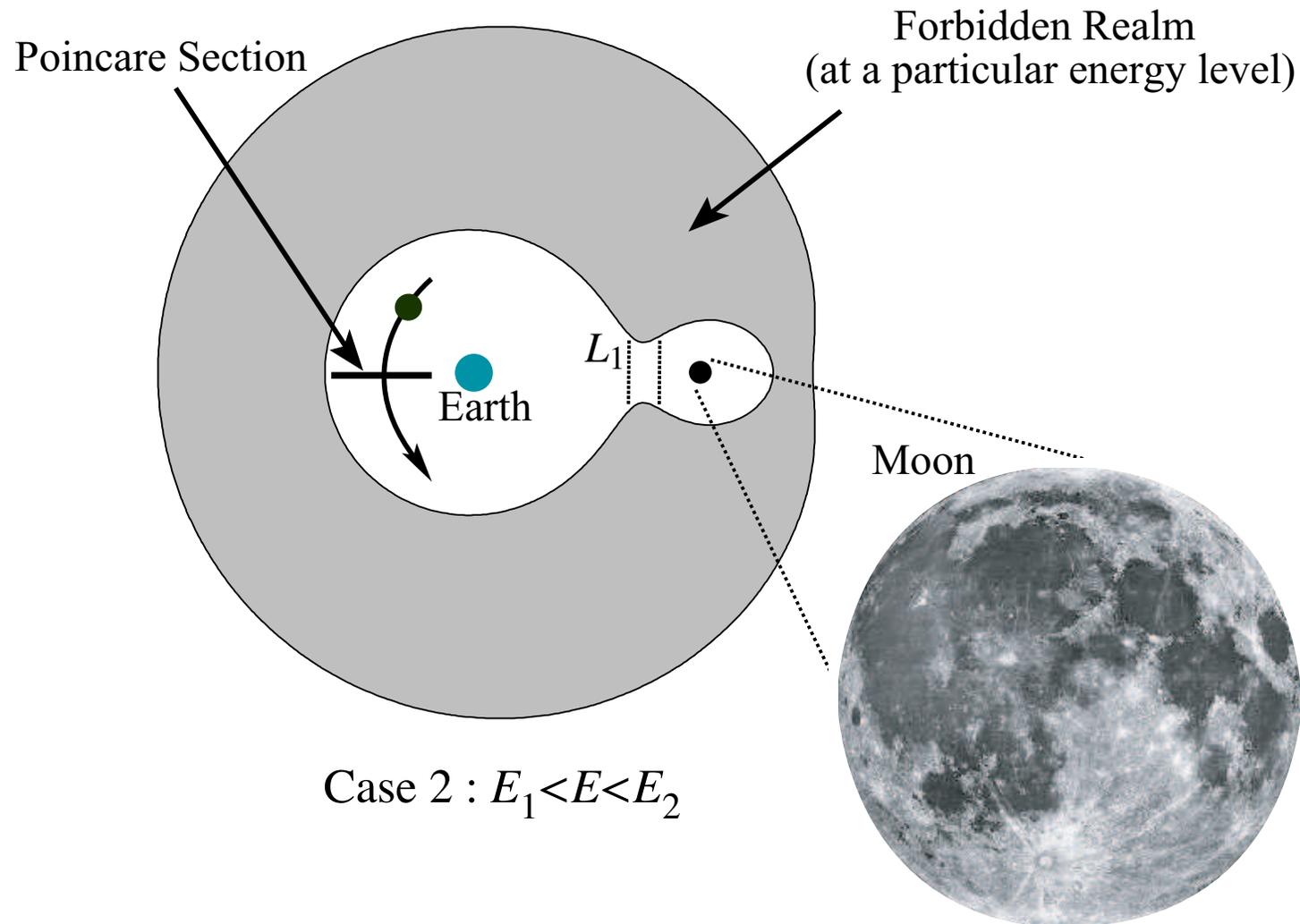
- Take full advantage of all known phase space structures: seek intersections between **resonances** and **tubes** leading to ballistic capture by the Moon

■ *Building blocks*

- **Appropriate energy:** transfer at three-body energy where resonances and tube dynamics are important
- **Poincare section:** reduces problem to 2D
- **Resonant gravity assists:** maximize change in orbit during every lunar encounter
- **Tube dynamics:** get captured by the Moon

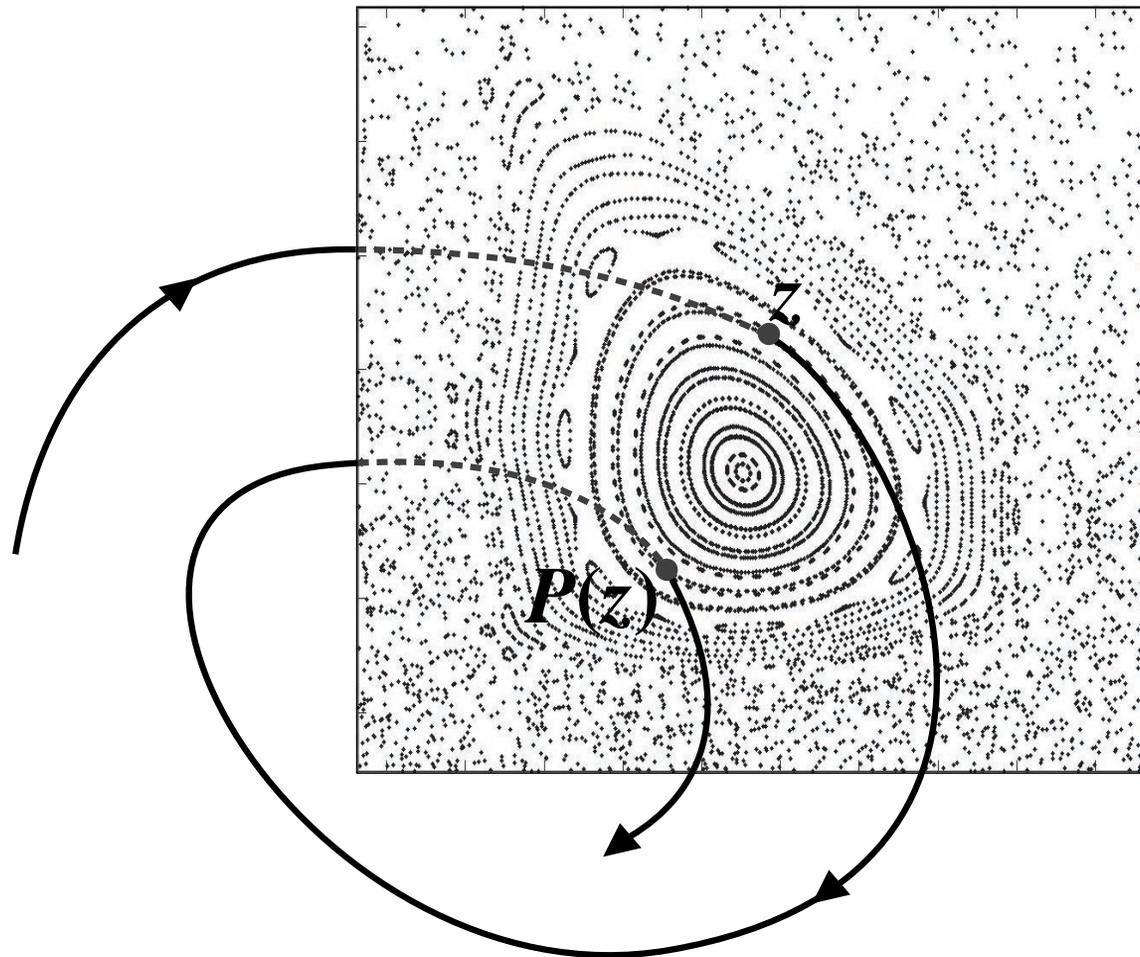
Appropriate Energy

- Transfer to occur on a single 3D energy surface
- Poincaré surface-of-section: motion on 2D map



Poincaré Surface of Section

- Study Poincaré surface of section at fixed energy E , reducing system to a 2-dimensional area preserving map.

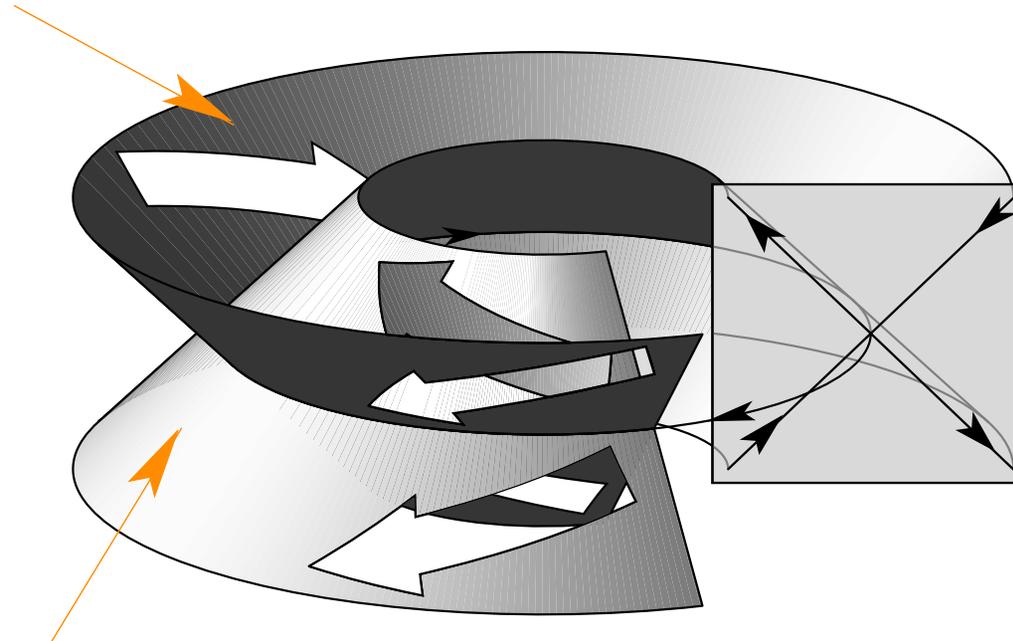


Poincaré surface of section

Resonant Gravity Assists

- Unstable resonances: Periodic orbits forming a dynamical “back-bone,” via their stable/unstable manifolds.
- Physically, these manifolds correspond to orbits undergoing repeated close encounters with the Moon.

Stable Manifold (orbits move toward the periodic orbit)

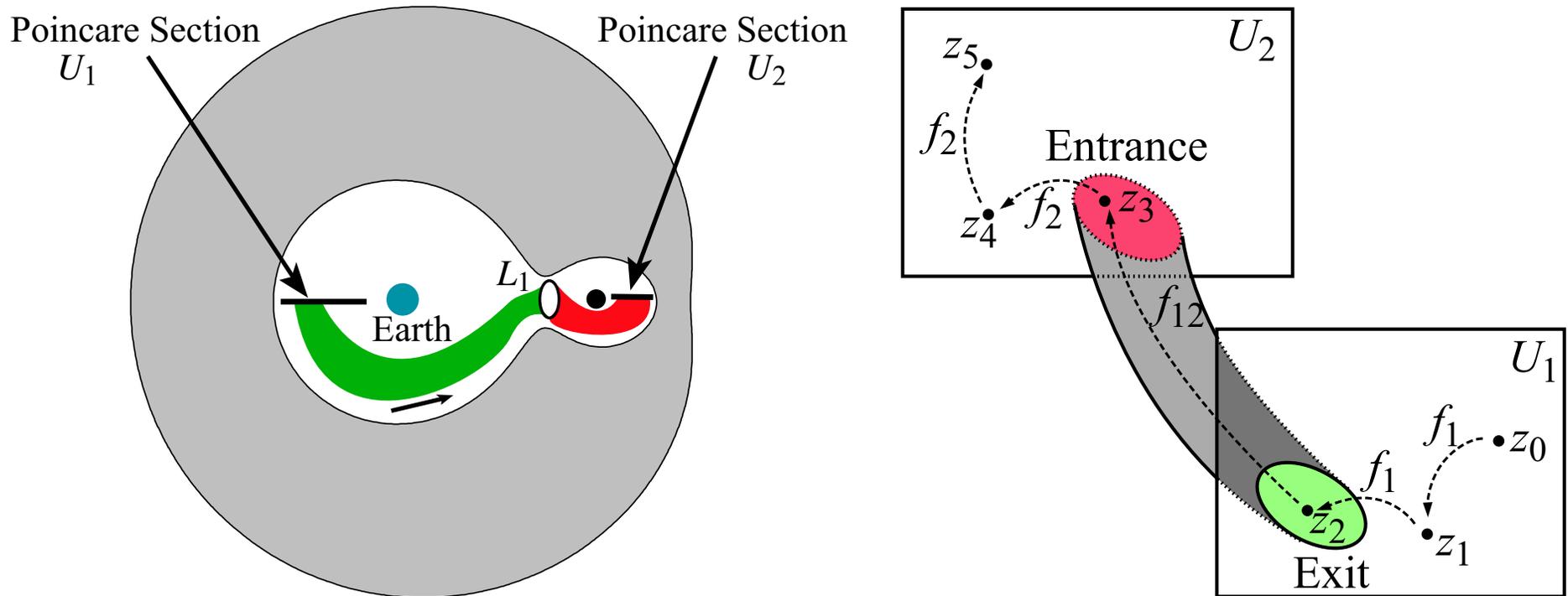


Unstable Manifold (orbits move away from the periodic orbit)

Unstable resonances and their manifolds.

Tube Dynamics

- Jump to the Moon's vicinity via invariant manifold tubes associated to a periodic orbit about L_1 .
- Track orbits via **exits/entrances** on Poincaré sections.



Tube dynamics: going from one Poincaré section to another.

Resonances and Tubes

■ *Resonances and tubes are linked*

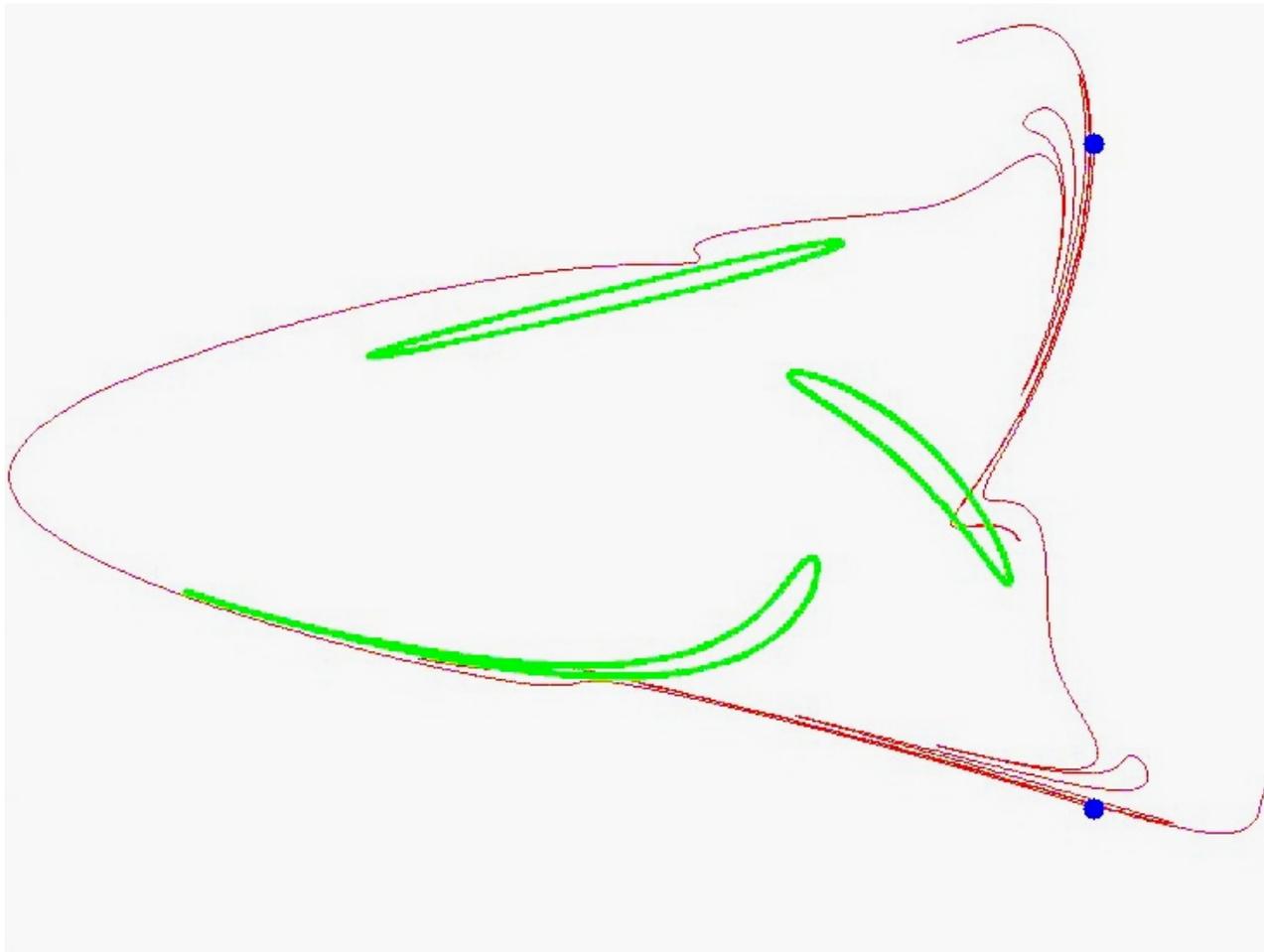
- It has been observed that the tubes of capture orbits are coming from certain resonances.
 - Koon, Lo, Marsden, Ross [2001]

■ *Designing an efficient transfer*

- First, from the starting Earth orbit, perform a ΔV placing the spacecraft on a trajectory near one of the resonances which is linked to capture tubes.
- Then perform small maneuvers to steer into a capture tube – none may be needed!

Resonances and Tubes

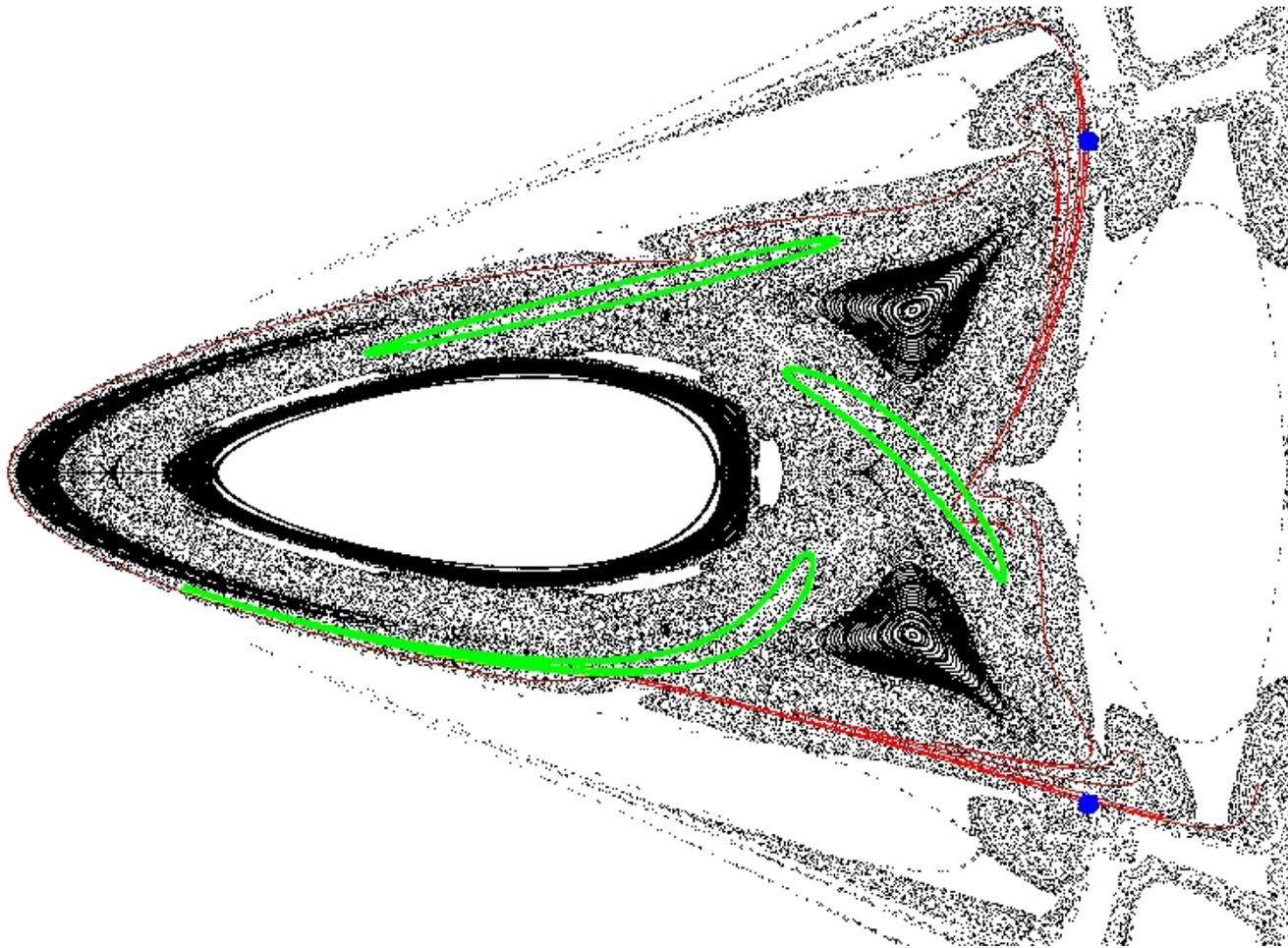
- Poincaré section: tube cross-sections are closed curves and resonance manifolds are windy curves.



Poincaré section showing tubes and resonances.

Resonances and Tubes

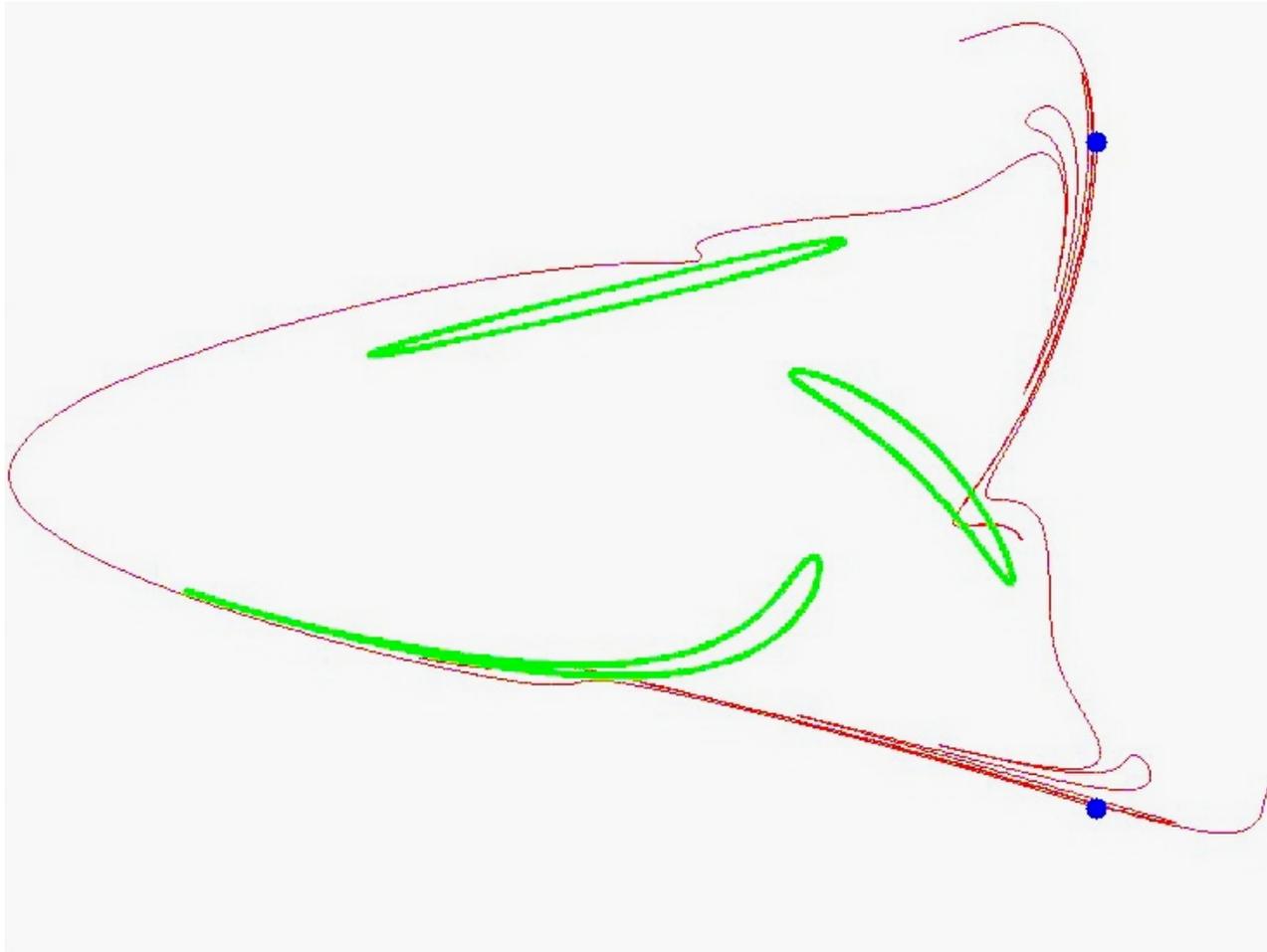
- Invisible structure: both of them reveal the structure in the “chaotic” part of phase space.



Poincaré section showing tubes and resonances and background points.

Resonances and Tubes

- Tubes and resonance manifolds intersect, i.e., there is a free transfer from the resonance to the Moon's vicinity.



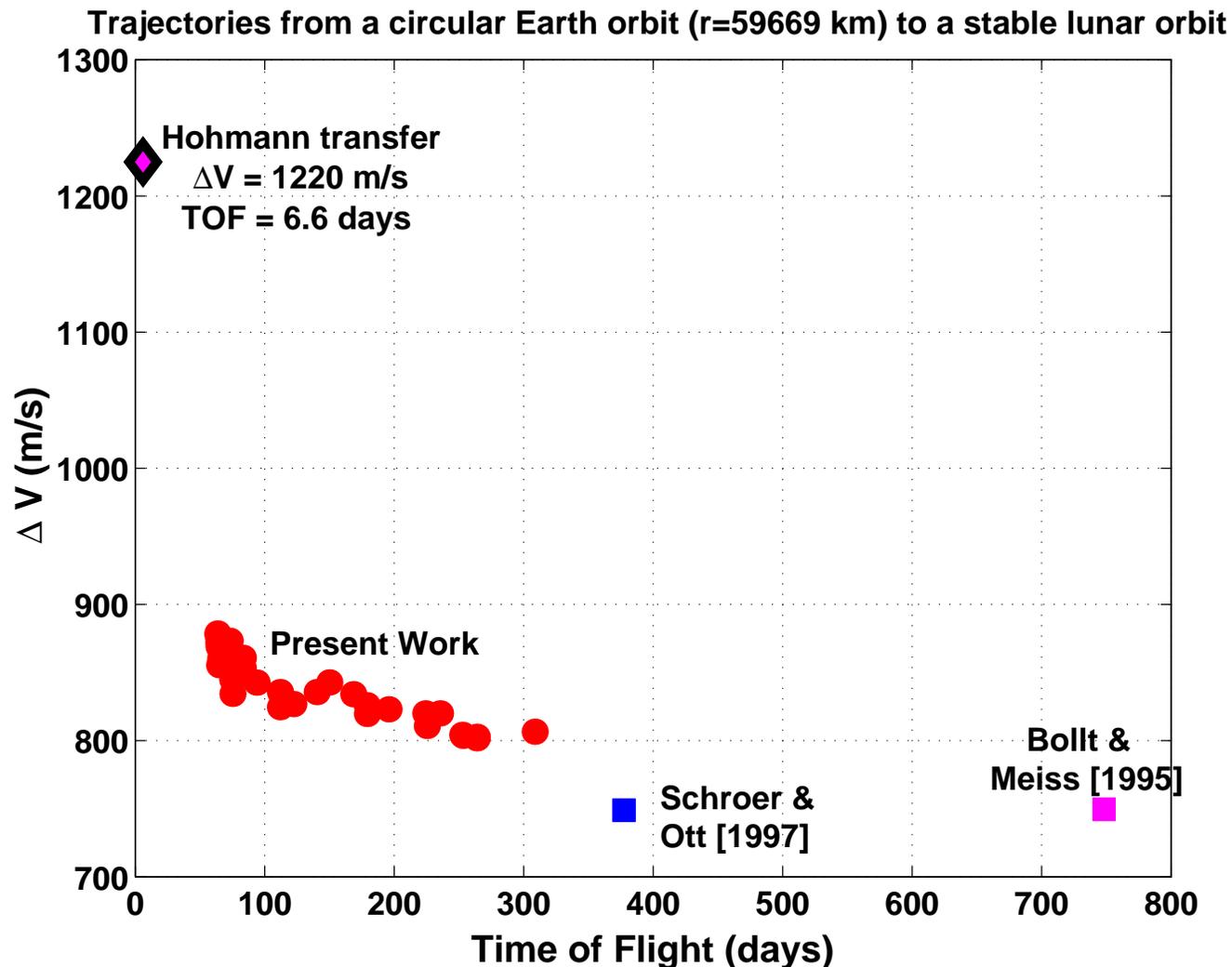
Poincaré section showing tubes and resonances.

Resonances and Tubes

- The free transfer corresponds to a flight time of over 250 days, but we can do better.
- The tubes and resonance manifold come “close” to intersecting in several places, i.e., a small ΔV can drastically cut out unnecessary flight time.
- By searching for ΔV 's of a particular size, one can systematically get a curve of ΔV vs. time of flight.

Resonances and Tubes

- **Results:** much shorter transfer times than previous authors for only slightly more ΔV

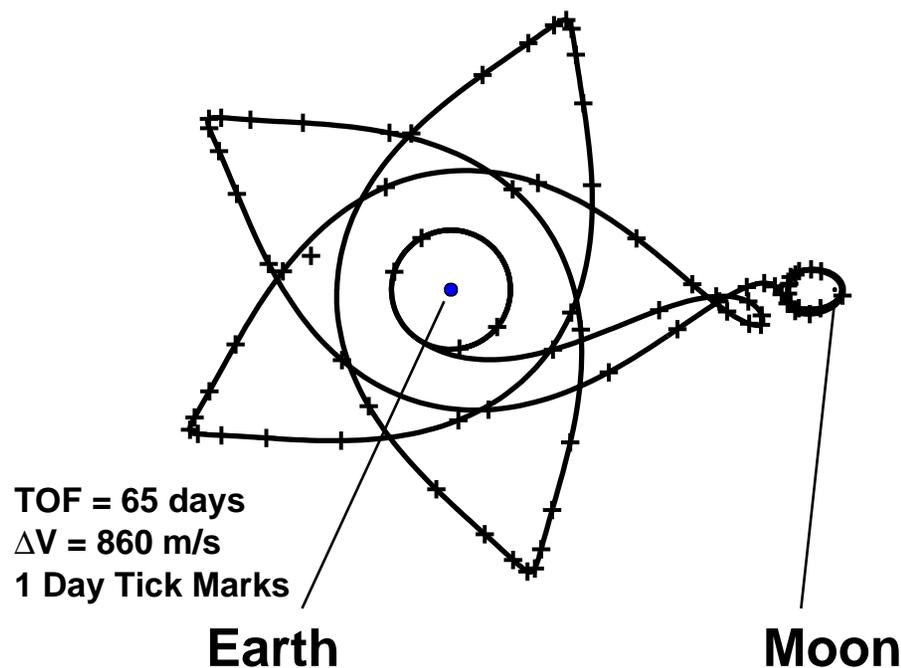


Resonances and Tubes

- Compare with Boltt and Meiss [1995]
 - A tenth of the time for only 100 m/s more

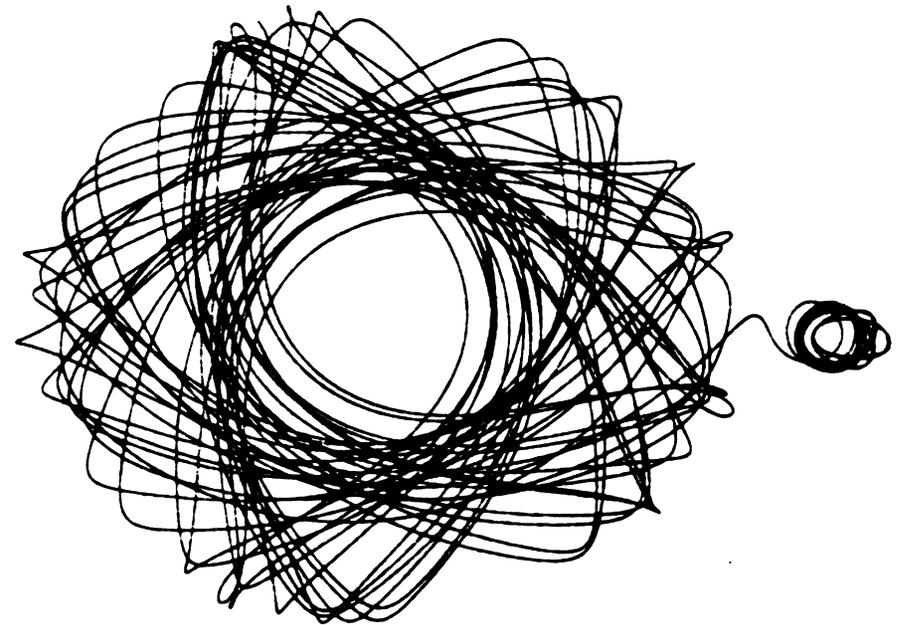
Current Result

65 days, $\Delta V = 860$ m/s



Boltt and Meiss [1995]

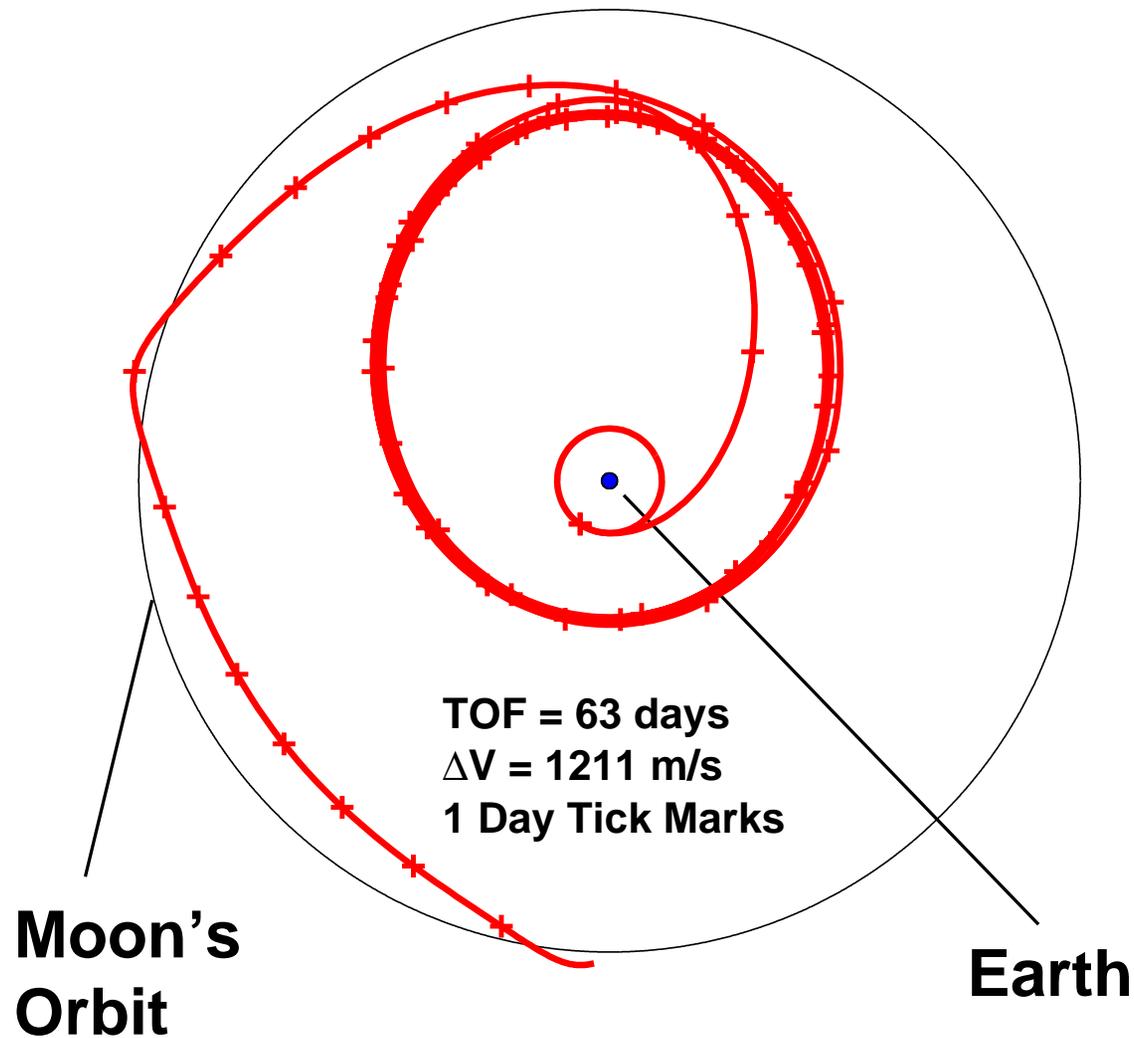
748 days, $\Delta V = 750$ m/s



e.g., GEO to Lunar Orbit

GEO to Moon Orbit Transfer

Seen in Geocentric Inertial Frame



e.g., GEO to Lunar Orbit

GEO to Moon - rotating frame

e.g., GEO to Lunar Orbit

GEO to Moon - inertial frame

References

- *Trade-Off Between Fuel and Time Optimization*, in preparation.
- Koon, W.S., M.W. Lo, J.E. Marsden and S.D. Ross [2002] *Constructing a low energy transfer between Jovian moons*. *Contemporary Mathematics* 292, 129–145.
- Gómez, G., W.S. Koon, M.W. Lo, J.E. Marsden, J. Masdemont and S.D. Ross [2001] *Invariant manifolds, the spatial three-body problem and space mission design*. AAS/AIAA Astrodynamics Specialist Conference.
- Koon, W.S., M.W. Lo, J.E. Marsden & S.D. Ross [2001] *Resonance and capture of Jupiter comets*. *Celestial Mechanics & Dynamical Astronomy* 81(1-2), 27–38.
- Koon, W.S., M.W. Lo, J.E. Marsden and S.D. Ross [2000] *Heteroclinic connections between periodic orbits and resonance transitions in celestial mechanics*. *Chaos* 10(2), 427–469.

For papers, movies, etc., visit the websites:

<http://www.cds.caltech.edu/~marsden>

<http://www.cds.caltech.edu/~shane>