

Dynamical Systems and Space Mission Design

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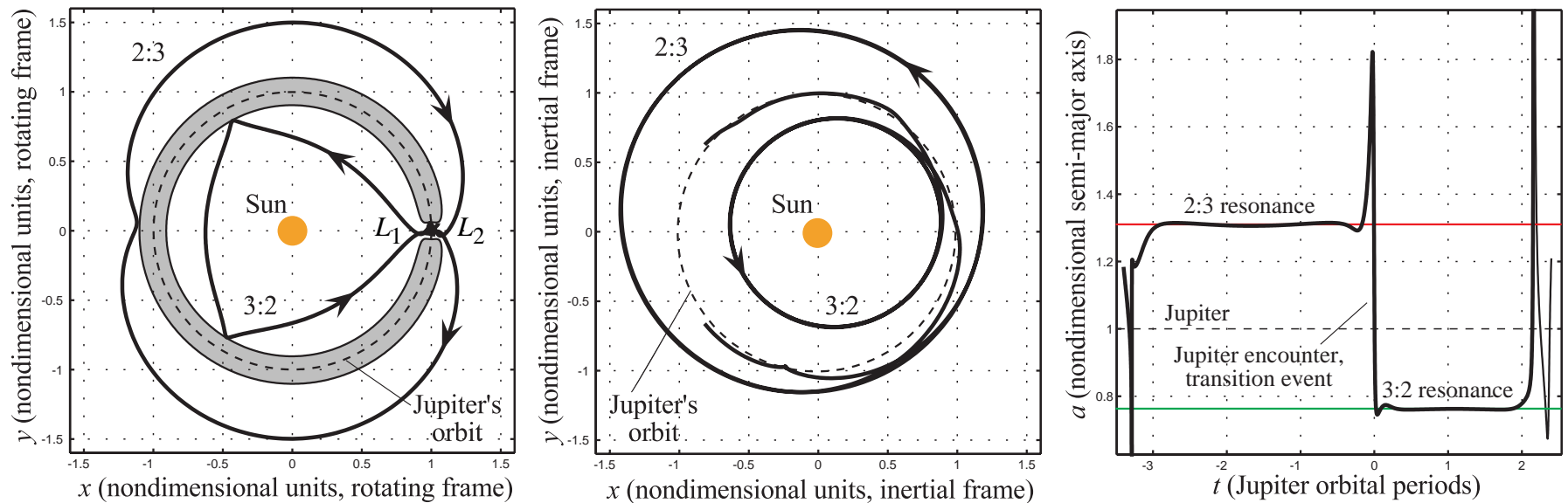
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■ Resonance Transition of Comets: Outline

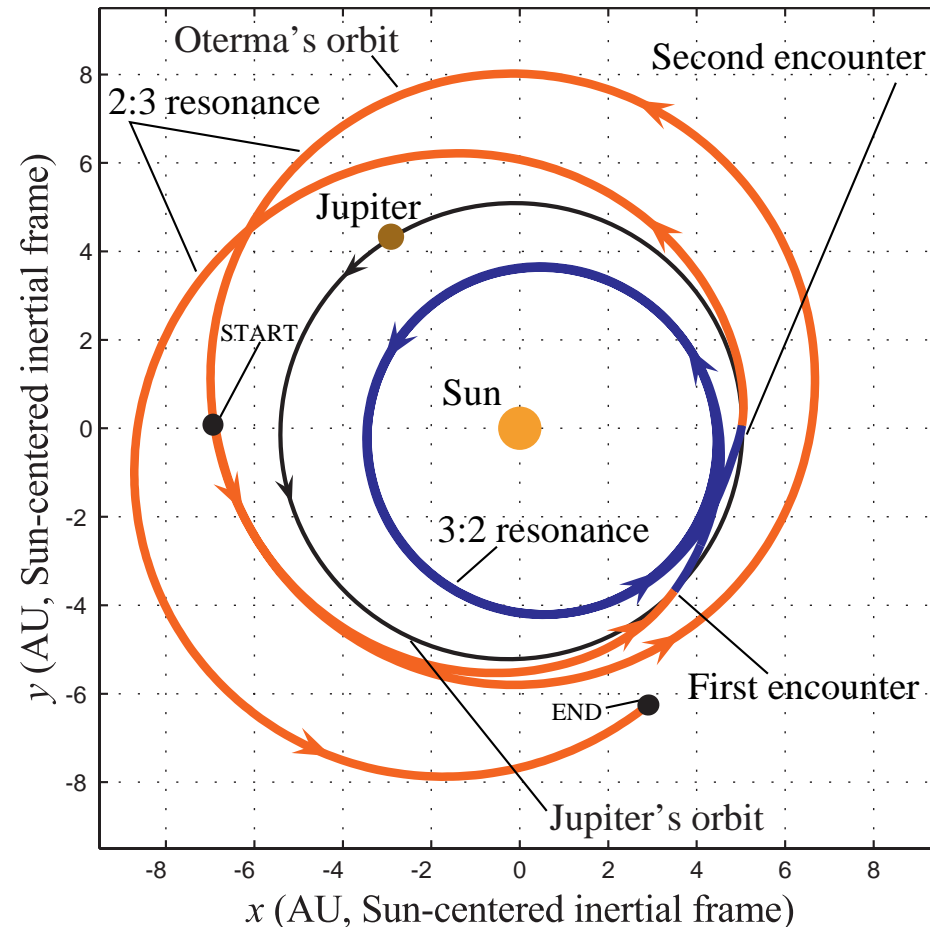
▶ Outline of Lecture 4B:

- **Resonance transition** seen in comets such as *Oterma*.
- Mixed phase space of 3-body problem:
- Mean motion resonance “**islands**” imbedded in chaotic “**sea.**”
- Exterior and interior resonances connected by Lyapunov orbit stable & unstable manifold tubes, the *dynamical channels*.
- **Future work:** transition between planets, belts, etc.



■ Jupiter Comets: *Oterma*

- ▶ Some Jupiter comets perform a **rapid transition** from the **outside** to the **inside** of Jupiter's orbit.
- ▶ **Captured temporarily** by Jupiter during transition.
- ▶ **Exterior** (**2:3 resonance**). **Interior** (**3:2 resonance**).

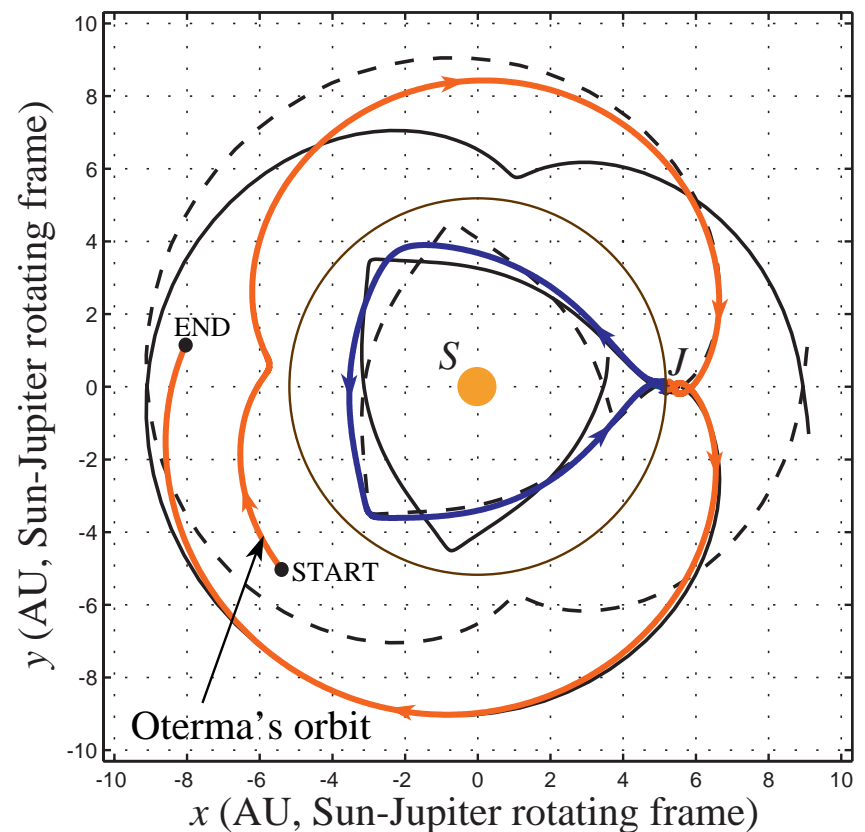
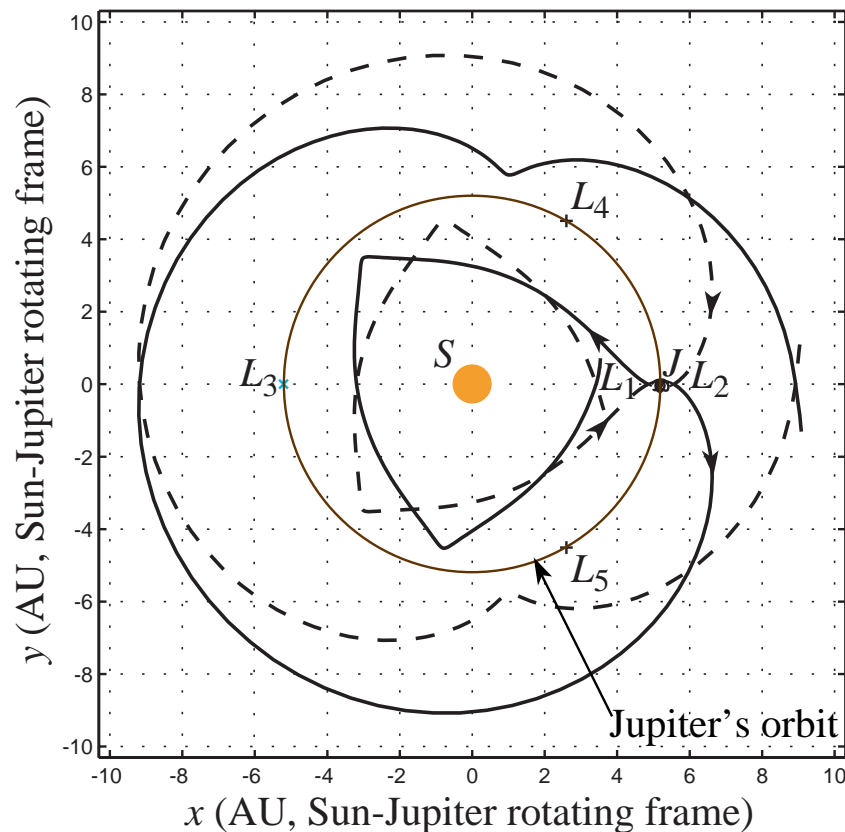


■ Jupiter Comets: Previous Works

▶ Belbruno/B. Marsden [1997]

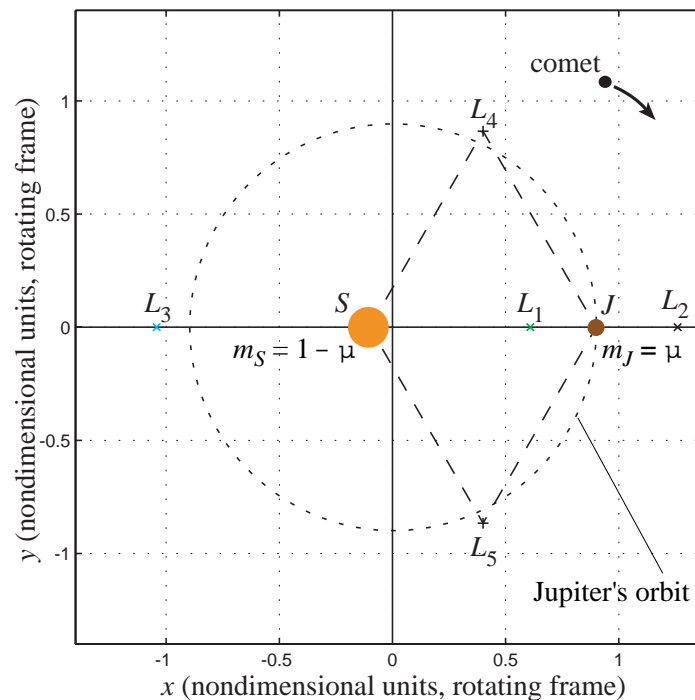
▶ Lo/Ross [1997] :

- Jupiter comets *Oterma*, *Gehrels 3*, etc. in Sun-Jupiter **rotating frame** follow **stable and unstable invariant manifolds** of the equilibrium points L_1 and L_2 .



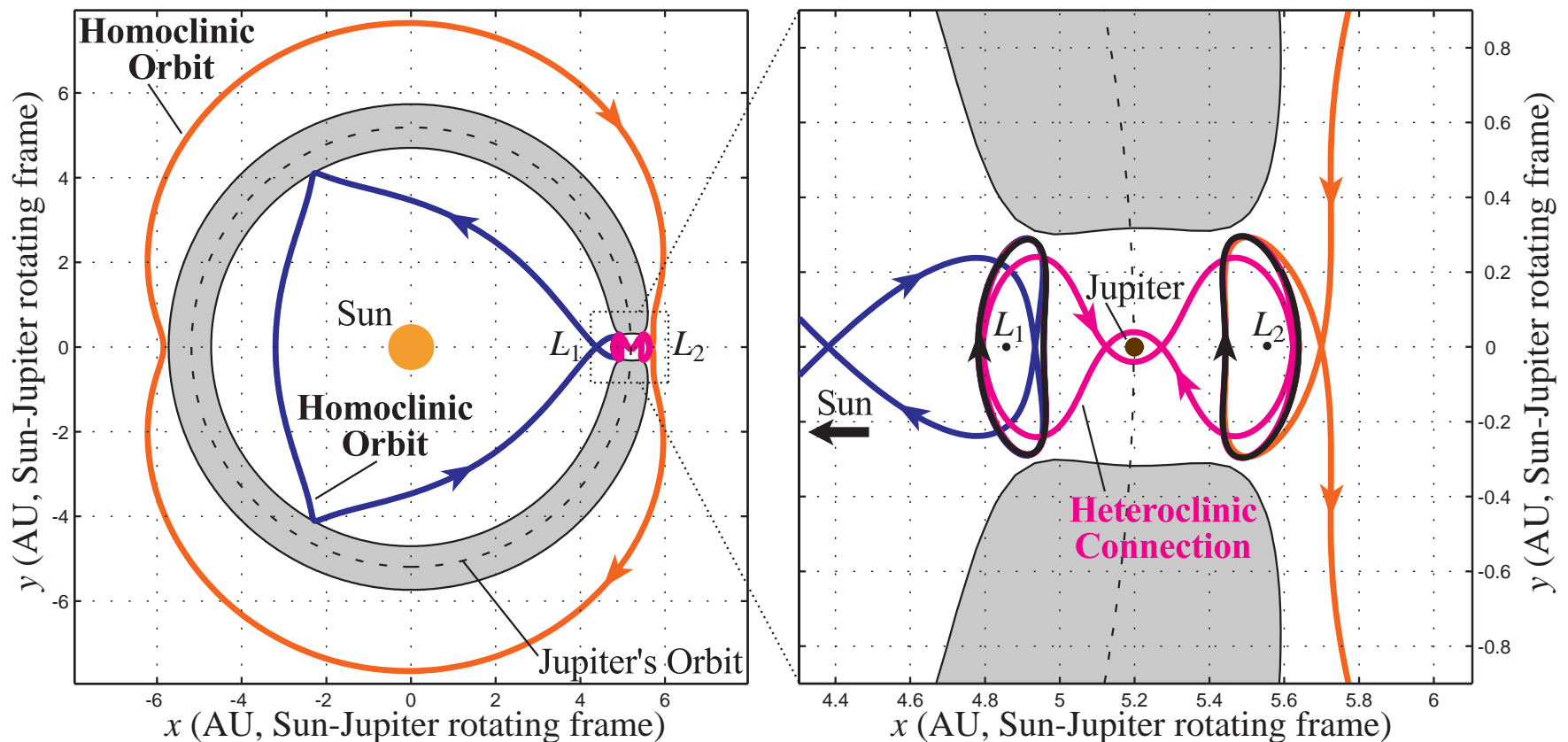
■ Jupiter Comets: Planar CR3BP Model

- ▶ Use planar circular restricted 3-body problem as initial model:
 - **Simplest 3-body model**, easiest to analyze.
 - Comets of interest are mostly **heliocentric**, but their perturbation is dominated by **Jupiter's gravitation**.
 - Their motion is nearly in Jupiter's **orbital plane** ($i < 5^\circ$), and Jupiter's small **eccentricity** (0.0483) plays little role during resonance transition.



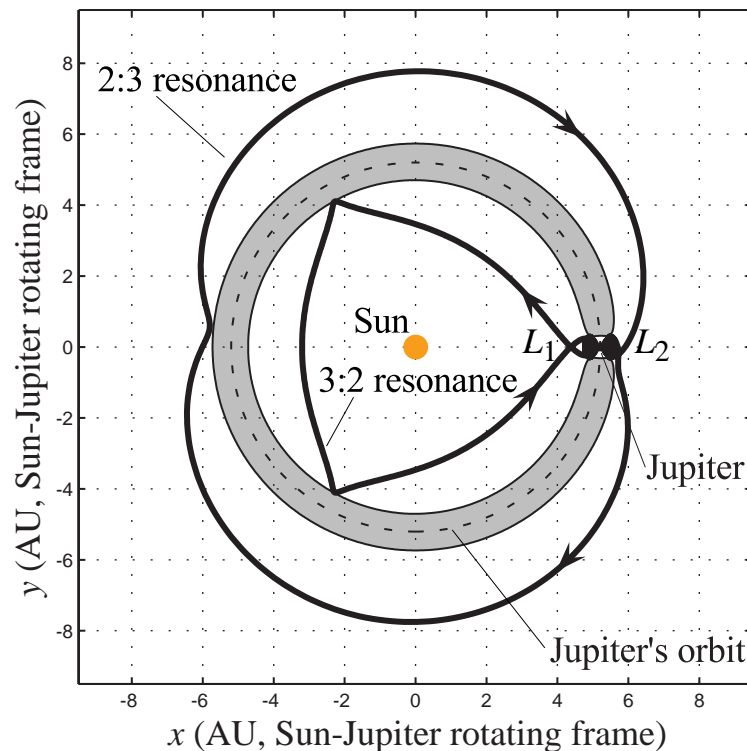
■ Jupiter Comets: Heteroclinic Connection

- ▶ More recently, Koon/Lo/Marsden/Ross [2000]:
 - Found **heteroclinic connection** between pair of periodic orbits.
 - Found a large class of **orbits** near this (homo/heteroclinic) **chain**.
 - Comets can follow these **dynamical channels** in rapid transition between **interior** and **exterior** Hill's regions.

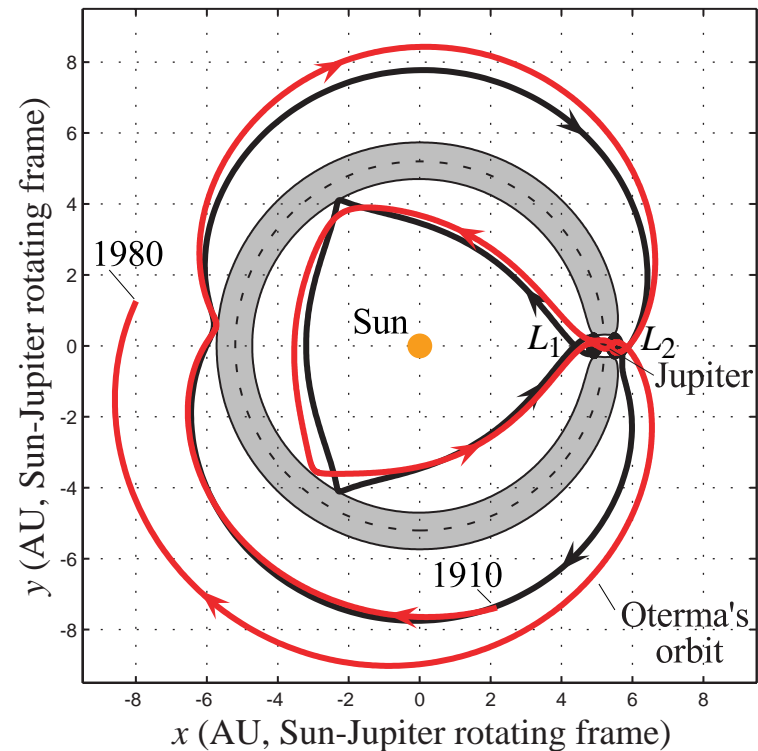


■ Jupiter Comets: Following Dynamical Channels

- For instance, consider the comet *Oterma* from 1910 to 1980.
- The average Jacobi constant for *Oterma* during its transition is $C = 3.030 \pm 0.005$ (computed at Jupiter encounter).
- We can compute a **homoclinic-heteroclinic chain** for $C = 3.030$ (shown in **black** on the left).
- Overlaying the chain, we plot *Oterma's* orbit in **red** (at right).



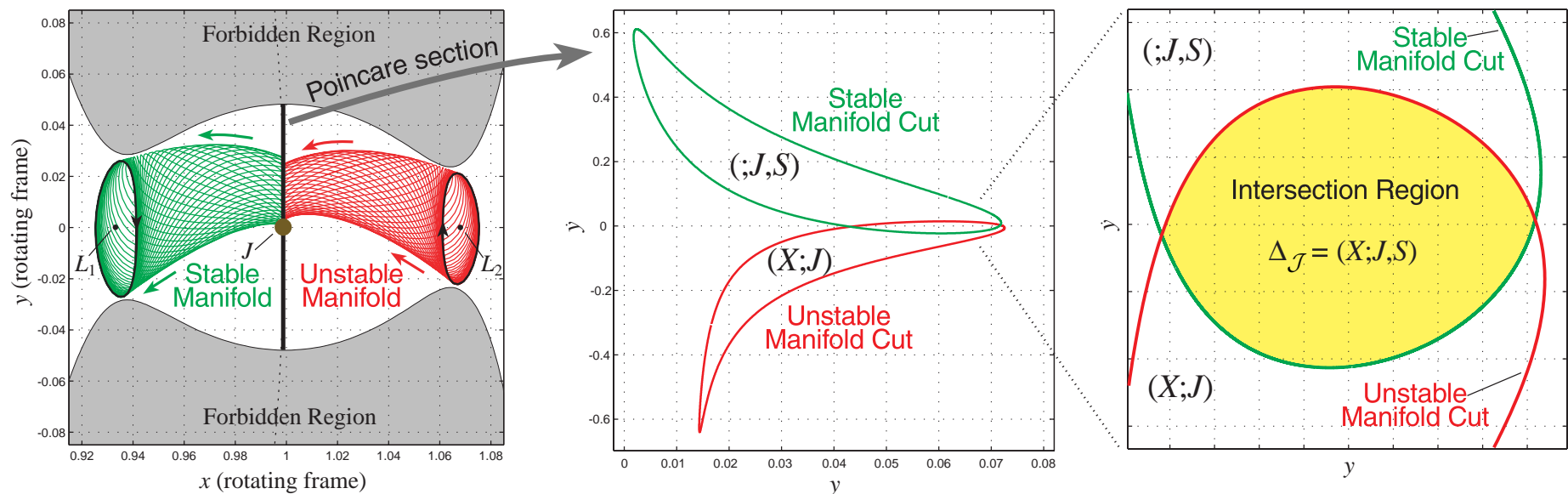
(a)



(b)

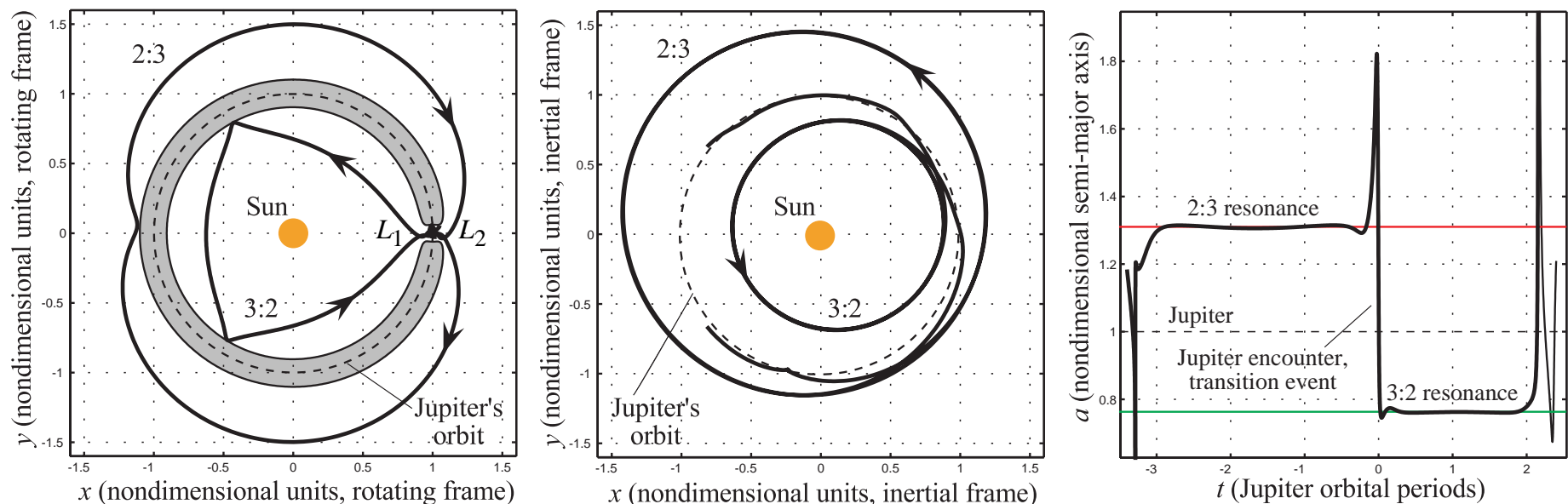
■ Jupiter Comets: Rapid Transition Mechanism

- **Rapid transition** between the interior and exterior regions is possible via the L_1 & L_2 periodic orbit **stable** & **unstable** manifold **tubes** (containing **transit** orbits) and their **intersections**.
- This was a surprising result. Some believed that a third degree of freedom was necessary for such a transition, or that “Arnold diffusion” was involved.
- But as we have seen, only the **planar CR3BP**, the simplest model of 3-body gravitational dynamics, is necessary.



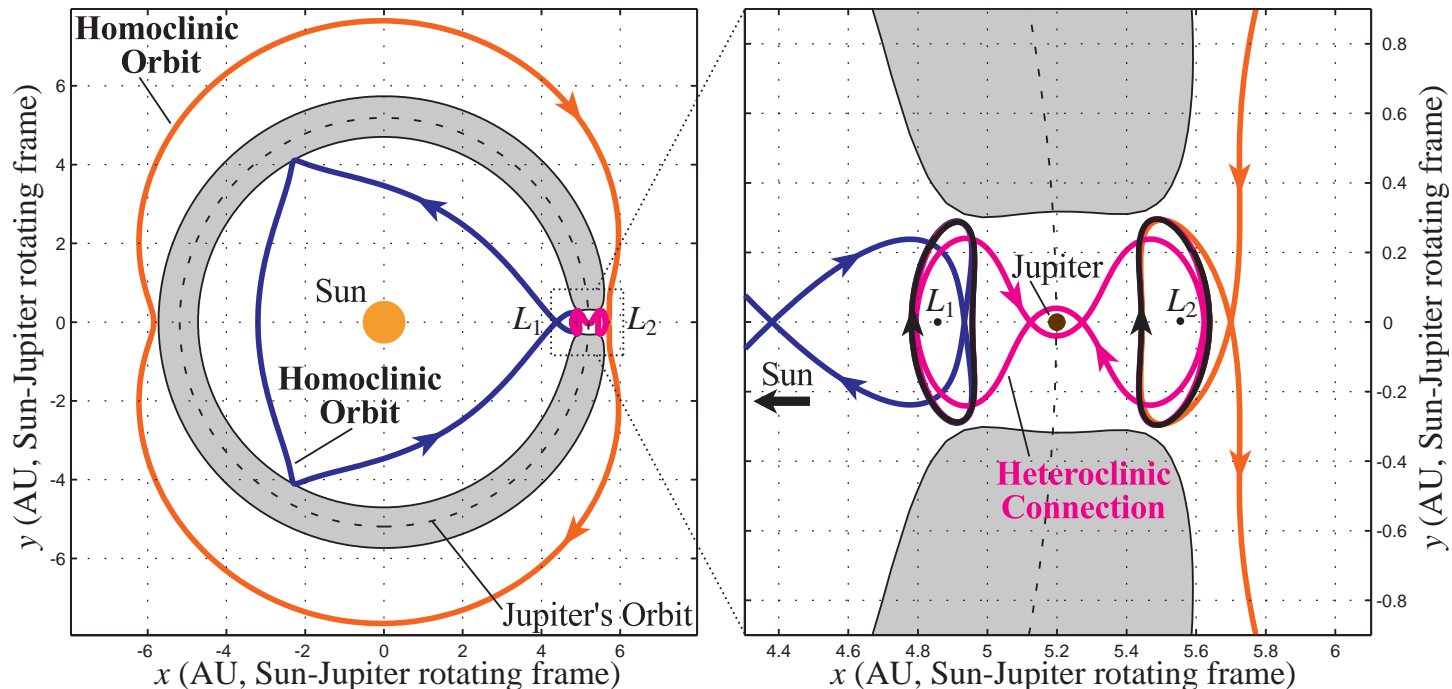
■ Rapid Transition Mechanism: Resonance Transition

- The tubes are a generic transport mechanism connecting the interior and exterior Hill's regions.
- We wish to understand their role in transport between interior and exterior **mean motion resonances**.
- e.g., we shall try to explain in more precise terms the sense in which *Oterma* transitions between the 3:2 and 2:3 mean motion resonances with Jupiter.



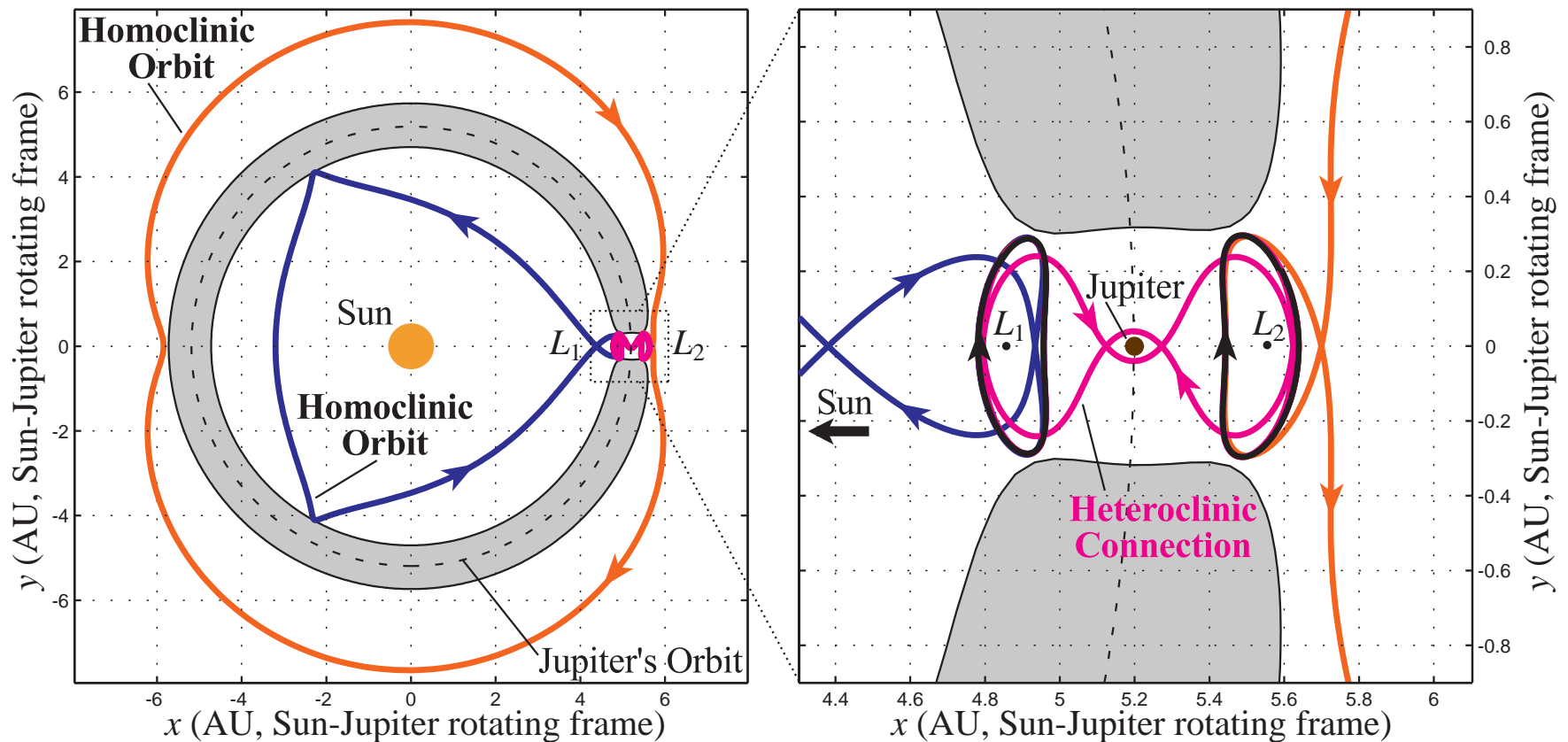
■ Rapid Transition Mechanism: Resonance Transition

- For the Sun-Jupiter system ($\mu = 0.0009537$), we can construct a homoclinic-heteroclinic chain with Jacobi constant similar to that of *Oterma* during its recent Jupiter encounters ($C = 3.030$).
- The chain is a union of orbits: **interior** region orbit homoclinic to L_1 periodic orbit, **exterior** region orbit homoclinic to L_2 periodic orbit, and **heteroclinic connection** between the L_1 & L_2 periodic orbits.



■ Rapid Transition Mechanism: Resonance Transition

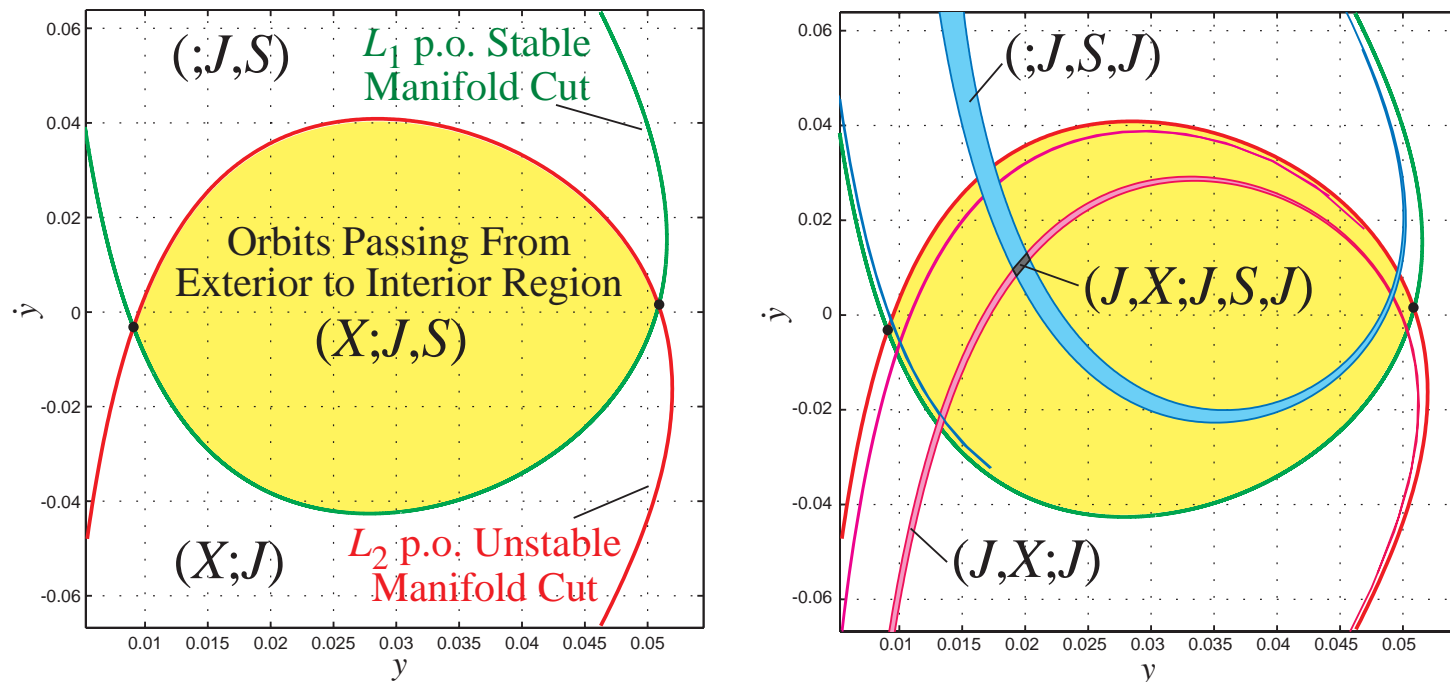
- We choose this chain because its homoclinic orbits are (1,1)-type.
- Limiting to (1,1)-type means, for this particular energy regime, that two different resonance connections are possible; 3:2 to 2:3, and 3:2 to 1:2. This will be explained later. We choose **3:2** to **2:3**, since this matches *Oterma's* orbit.



■ Rapid Transition Mechanism: Resonance Transition

- Our *main theorem* tells us that in the vicinity of this chain, there exists an orbit whose symbolic sequence $(\dots, J, X, J, S, J, \dots)$ is periodic and has the *central block itinerary* (J, X, J, S, J) .
- This orbit transitions between the interior and exterior regions (the neighborhood of the 3:2 and 2:3 resonances, in particular). We call this kind of itinerary a *resonance transition block*.

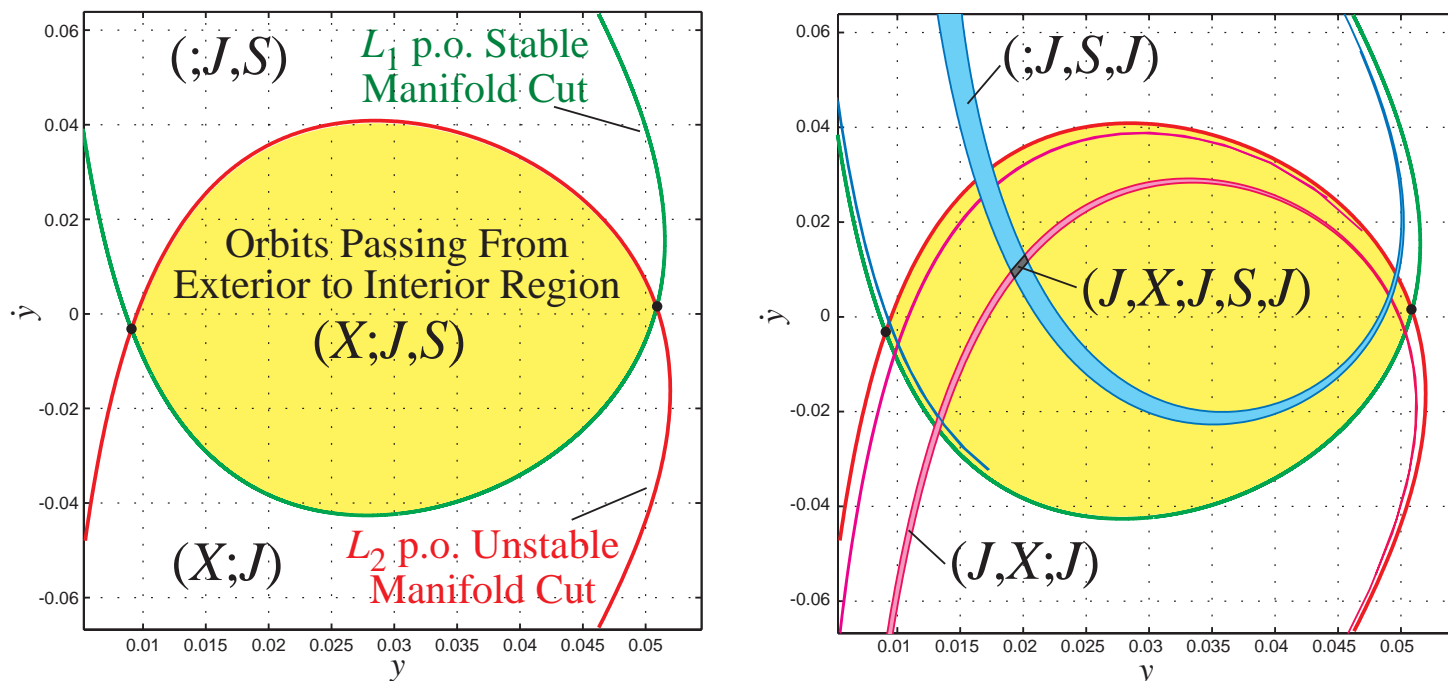
Poincare Section in the Jupiter Region



■ Rapid Transition Mechanism: Resonance Transition

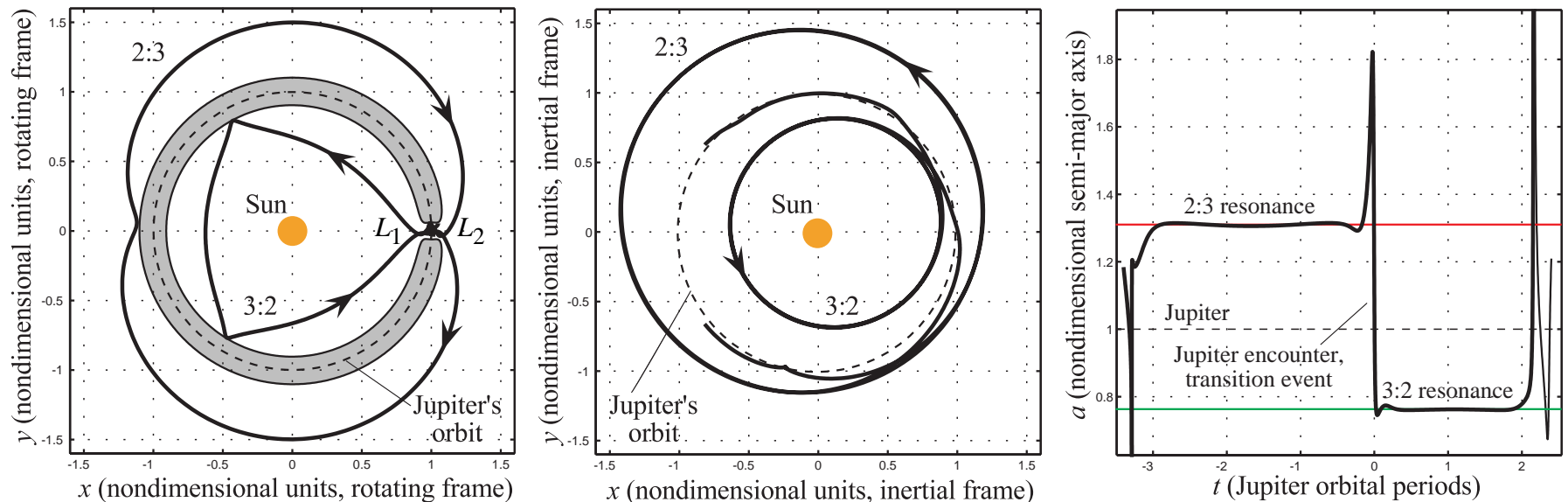
- This orbit makes a **rapid transition** from the exterior to the interior region and vice versa, passing through the Jupiter region. It will repeat this pattern **eternally**.
- While an orbit with this exact itinerary is very fragile, the structure of nearby orbits whose symbolic sequences have the same central block itinerary, namely (J, X, J, S, J) , is quite **robust**.

Poincare Section in the Jupiter Region



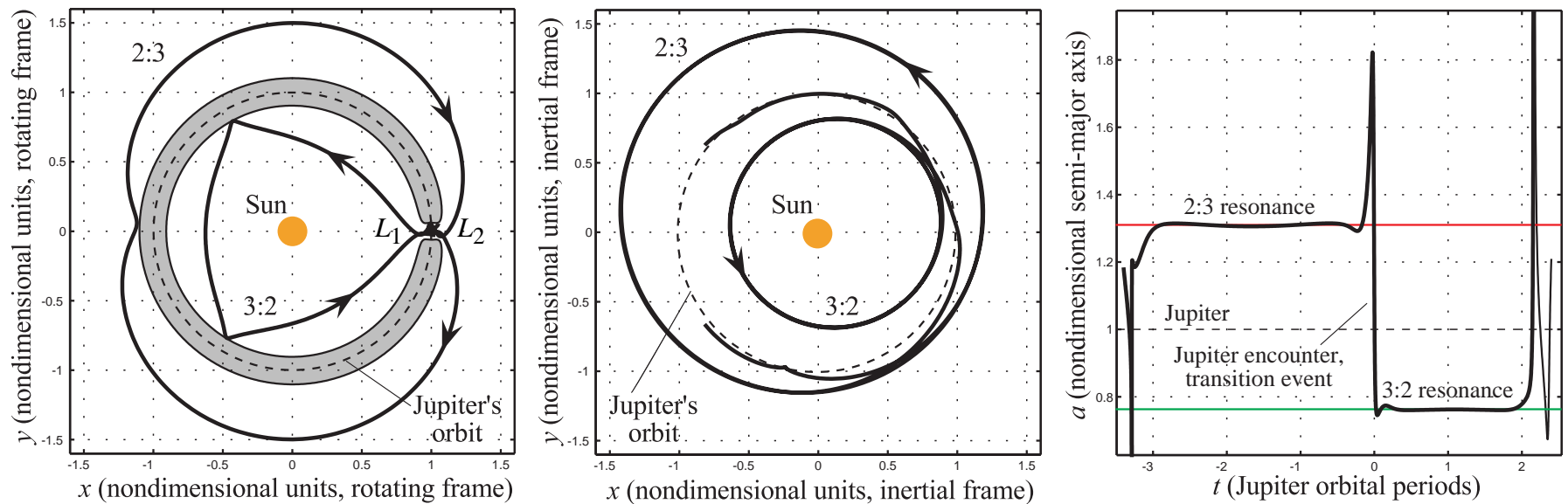
Rapid Transition Mechanism: Resonance Transition

- An example orbit with central block (J, X, J, S, J) is shown below.
- We will study how the dynamical channels near the chain connect the **3:2 resonance** of the interior region with the **2:3 resonance** of the exterior region.



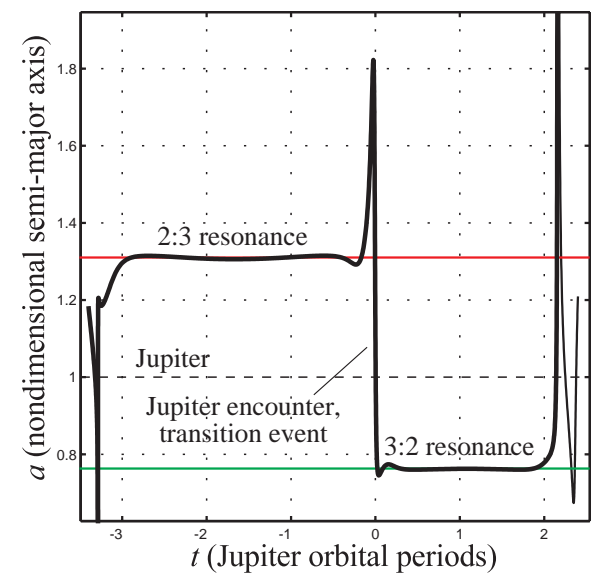
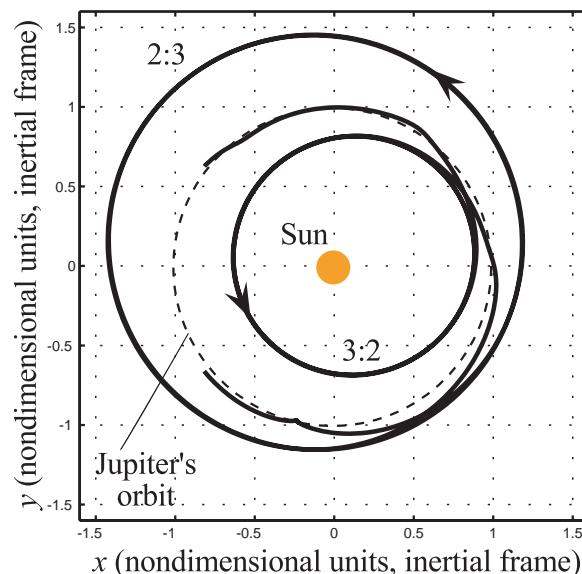
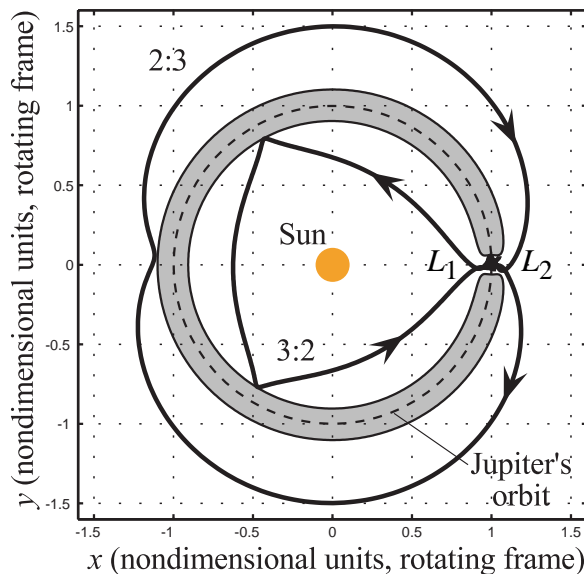
PCR3BP: Perturbation of the Two-Body Problem

- Recall that the PCR3BP is a *perturbation* of the two-body problem. Hence, outside of a small neighborhood of L_1 , the trajectory of a comet in the interior region follows essentially a *two-body orbit* around the Sun.
- In the heliocentric inertial frame, the orbit is nearly *elliptical*.



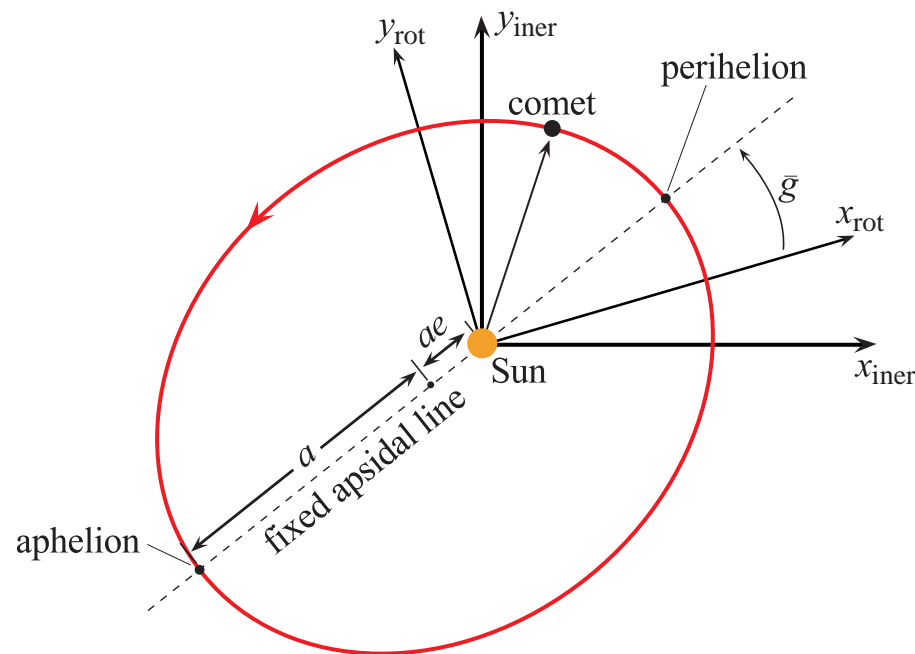
■ Heliocentric Orbits: Mean Motion Resonance

- The *mean motion resonance* of the comet with respect to Jupiter is equal to $a^{-3/2}$ where a is the semi-major axis of the heliocentric elliptical orbit. Recall that the Sun-Jupiter distance is normalized to be 1 in the PCR3BP.
- The comet is said to be in $p:q$ *resonance with Jupiter* if $a^{-3/2} \approx p/q$, where p and q are small integers. In the heliocentric inertial frame, the comet makes roughly p *revolutions* around the Sun in q *Jupiter periods*.



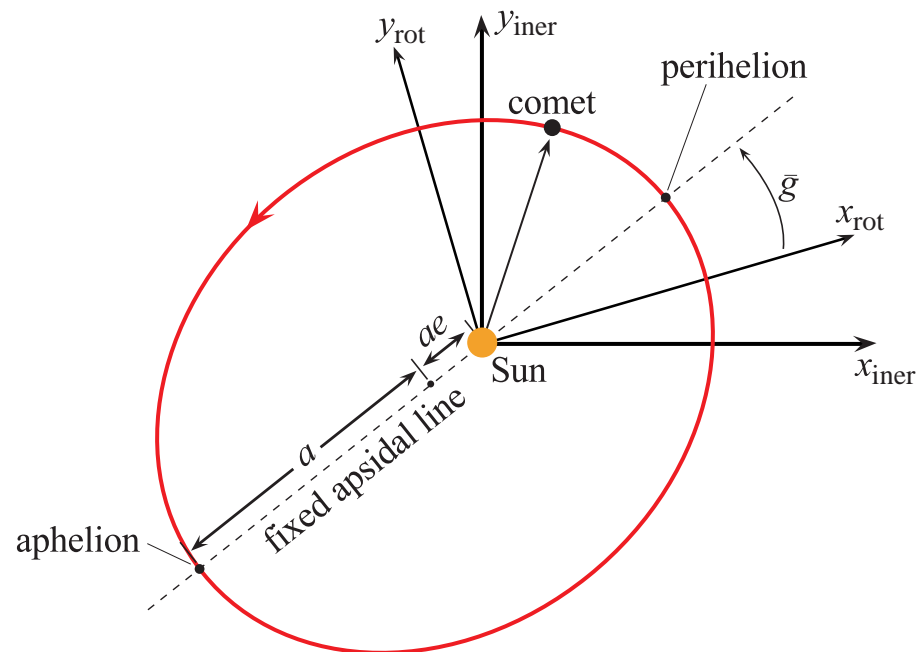
■ Canonical Coordinates: Delaunay Variables

- To study the process of resonance transition, we shall use a set of canonical coordinates, called ***Delaunay variables***, which make the study of the ***two-body regime*** of motion particularly simple.
- Delaunay variables in the rotating coordinates are denoted (l, \bar{g}, L, G) . $G = [a(1 - e^2)]^{1/2}$ is the angular momentum. L is related to the semi-major axis a , by $L = a^{1/2}$, hence encodes the ***mean motion resonance*** (with respect to Jupiter in the Sun-Jupiter system).

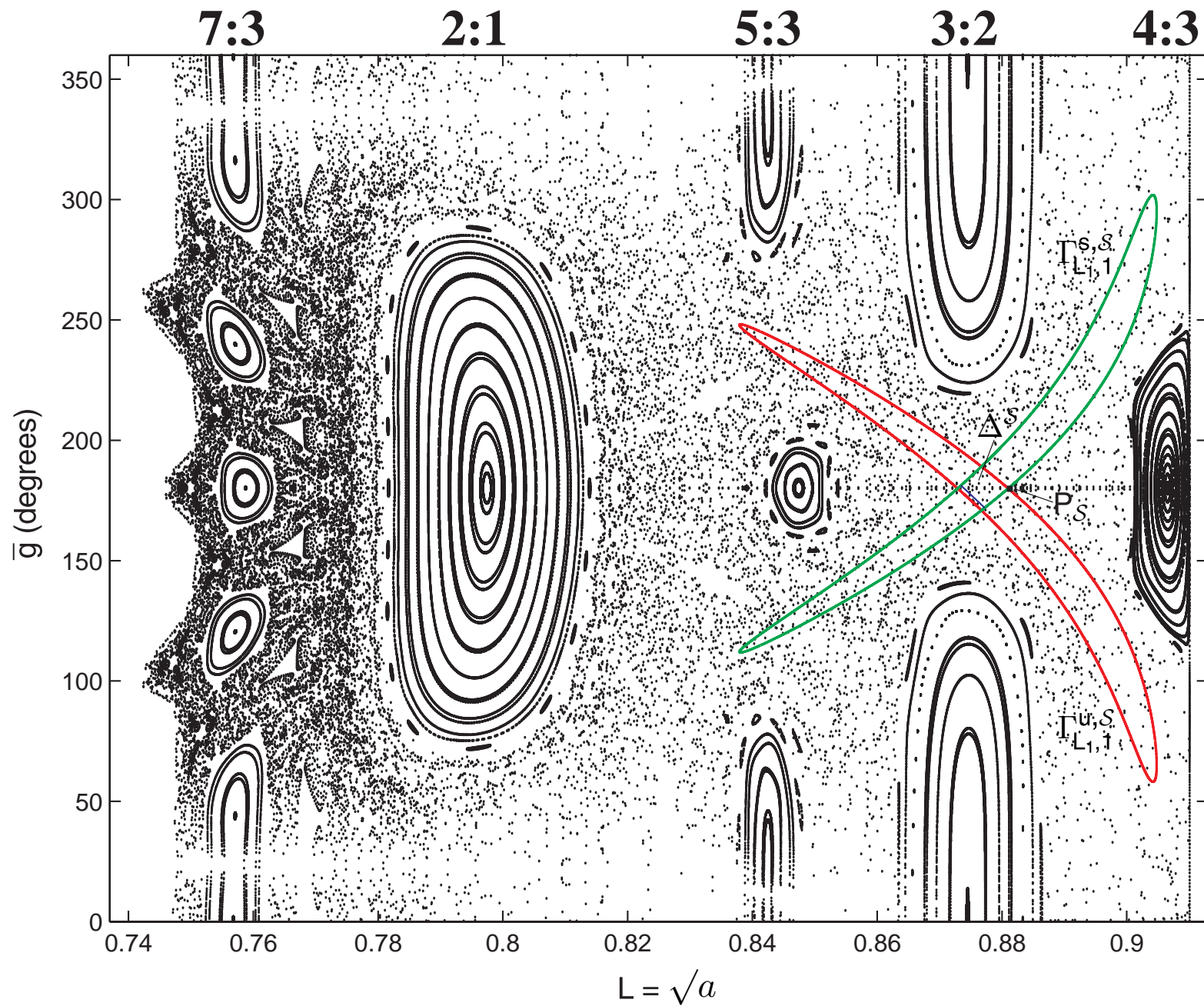


■ Canonical Coordinates: Delaunay Variables

- Both l and \bar{g} are angular variables defined modulo 2π .
- \bar{g} is the *argument of perihelion* relative to the rotating axis.
- l is the *mean anomaly*, the ratio of the area swept out by the ray from the Sun to the comet starting from its perihelion passage to the total area.
- Szebehely [1967], Abraham & Marsden [1978], Meyer & Hall [1992].



Interior Resonances: U_1 Poincaré Section (L, \bar{g})



■ Interior Resonances: U_1 Poincaré Section

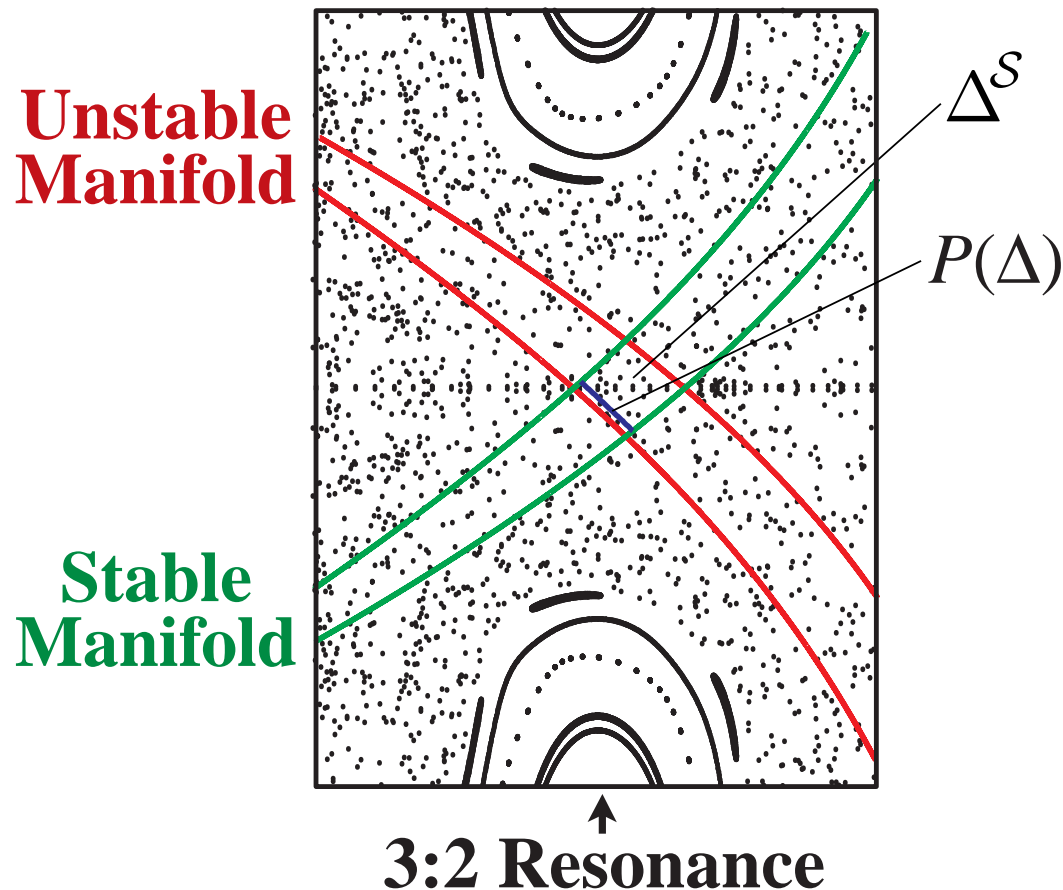
- The striking thing is that the first cuts of the **stable** and **unstable** manifolds *intersect* exactly at the region of the **3:2 resonance**.
- Their intersection Δ^S contains all the orbits that have come from the Jupiter region J into the interior region S , gone around the Sun once (in the rotating frame), and will return to the Jupiter region. In the heliocentric inertial frame, these orbits are nearly elliptical outside a neighborhood of L_1 .
- They have a semi-major axis which corresponds to 3:2 resonance by Kepler's law (i.e., $a^{-3/2} = L^{-3} \approx 3/2$). Therefore, *any* Jupiter comet which has an *energy similar to Oterma's* and which circles around the Sun *once* in the interior region *must be in 3:2 resonance* with Jupiter.

■ Mixed Phase Space: Stable “Islands” & Chaotic “Sea”

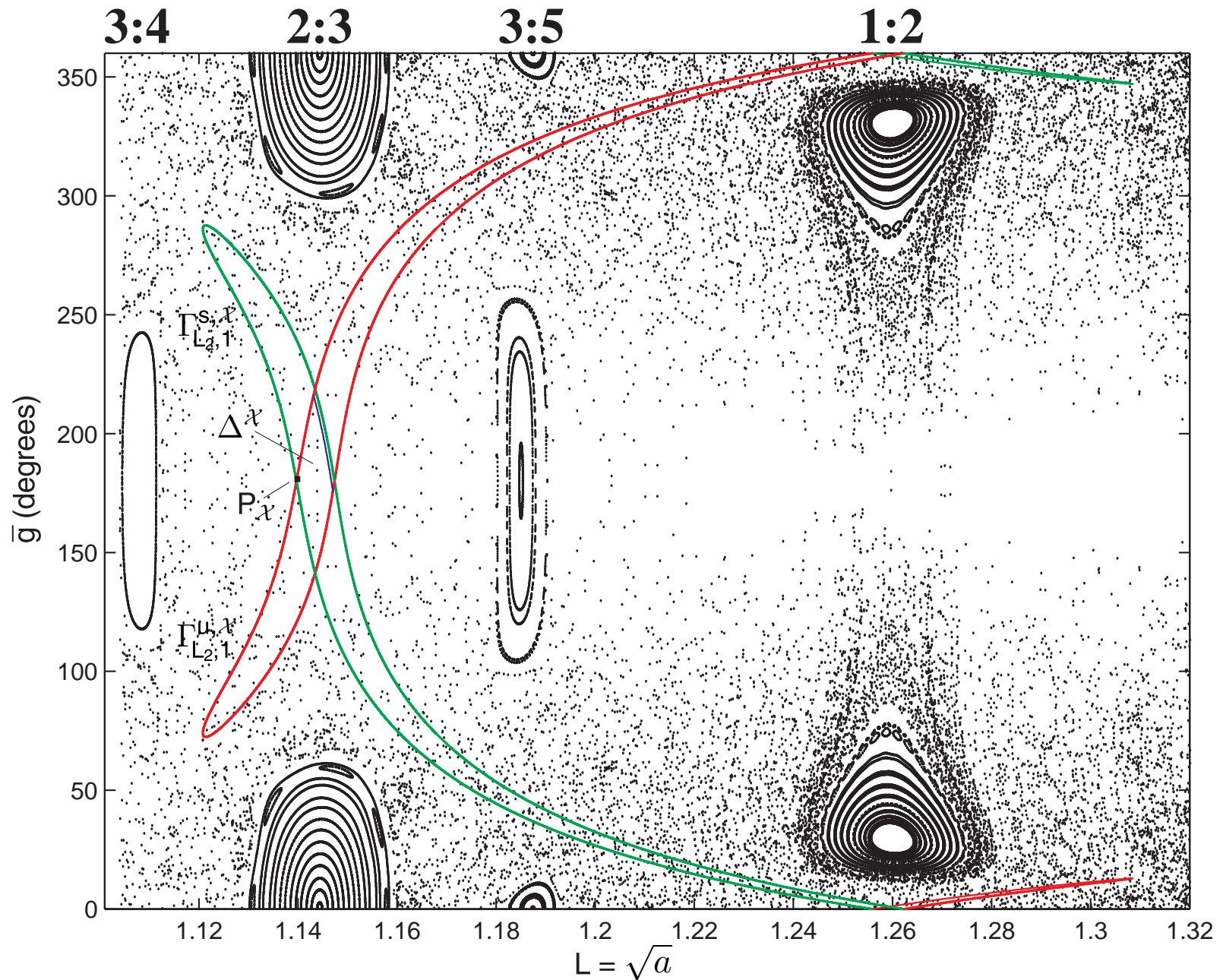
- The black background points on the U_1 Poincaré section reveal the character of the interior region phase space for this energy surface.
- ***Mixed phase space*** of stable periodic and quasiperiodic ***invariant tori “islands”*** embedded in bounded ***chaotic “sea.”***
- The ***families of stable tori***, where a “family” denotes those tori islands which lie along a strip of nearly constant L , correspond to ***mean motion resonances***. The size of the tori island corresponds to the dynamical significance of the resonance. The number of tori islands equals the order of the resonance (e.g., 3:2 is order 1, 5:3 is order 2).
- In the center of each island, there is a point corresponding to an exactly periodic, stable, resonant orbit. In between the stable islands of a particular resonance (i.e., along a strip of nearly constant L), there is a saddle point corresponding to an exactly periodic, unstable, resonant orbit. In the figure, the manifold intersection region Δ^S is centered on this saddle point for the 3:2 resonance.

■ Connection Between Interior and Exterior Resonances

- A subset of the interior resonance intersection region Δ^S is connected to exterior resonances through a heteroclinic intersection in the Jupiter region. This small *blue strip* inside Δ^S is part of the dynamical channel connecting interior and exterior resonances, and is thus the *resonance transition mechanism* we seek.



Exterior Resonances: U_4 Poincaré Section (L, \bar{g})



■ Exterior Resonances: U_4 Poincaré Section (L, \bar{g})

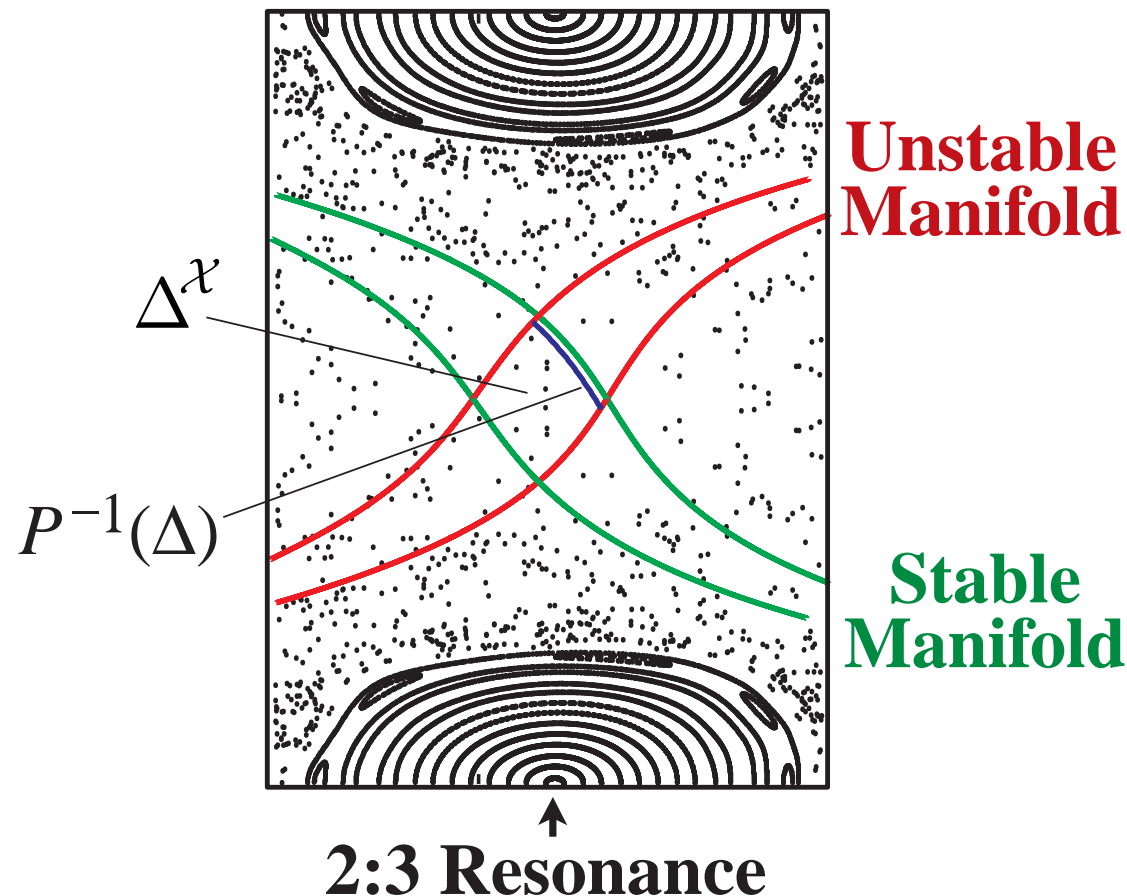
- We show the first exterior region Poincaré cuts of the **stable** and **unstable** manifolds of an L_2 periodic orbit with the U_4 section on the same energy surface.
- Notice that the first cuts of the **stable** and **unstable** manifolds intersect at *two places*; one of the intersections is exactly at the region of the *2:3 resonance*, the other is at the *1:2 resonance*.
- The intersection $\Delta^{\mathcal{X}}$ contains all the orbits that have come from the Jupiter region J into the exterior region X , have gone around the Sun once (in the rotating frame), and will return to the Jupiter region. Note that $\Delta^{\mathcal{X}}$ has two components (the 2:3 and 1:2 resonance regions).

■ Exterior Resonances: U_4 Poincaré Section (L, \bar{g})

- In the heliocentric inertial frame, these orbits are nearly elliptical outside a neighborhood of L_2 . They have a semi-major axis which corresponds to either 2:3 or 1:2 resonance by Kepler's law. Therefore, any Jupiter comet which has an energy similar to *Oterma's* and which circles around the Sun once in the exterior region must be in either 2:3 or 1:2 resonance with Jupiter.
- The background points were generated by a technique similar to those in the interior resonance Poincaré section. They reveal a similar mixed phase space, but now the resonances are exterior resonances (exterior to the orbit of Jupiter). We see that the exterior resonance intersection region $\Delta^{\mathcal{X}}$ envelops both the 2:3 and the 1:2 unstable resonance points.

■ Connection Between Interior and Exterior Resonances

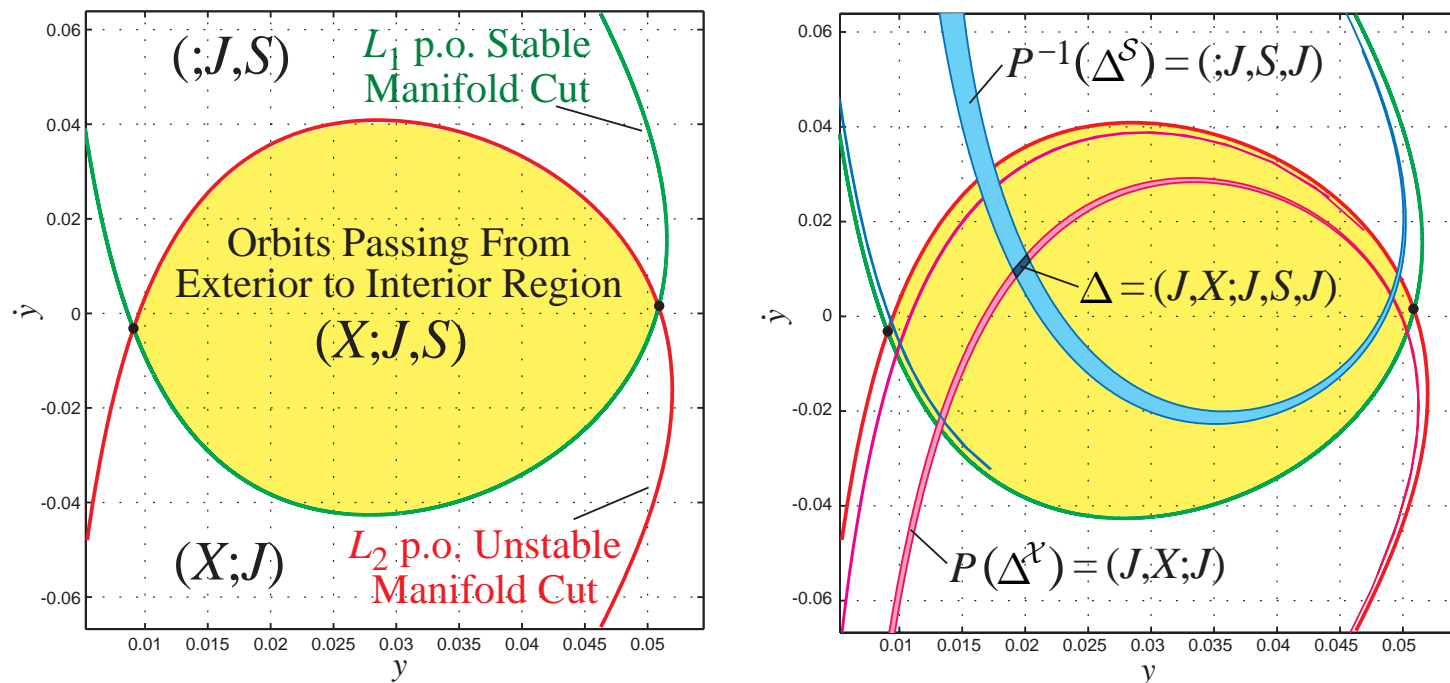
- A portion of $\Delta^{\mathcal{X}}$ is connected to interior resonances through a heteroclinic intersection in the Jupiter region. In particular, the small **blue strip** inside the **2:3 intersection region** connects to the **3:2 intersection region** of $\Delta^{\mathcal{S}}$ (and is its pre-image). We have thus found the **resonance transition** used by **Oterma**.



■ Connection Between Interior and Exterior Resonances

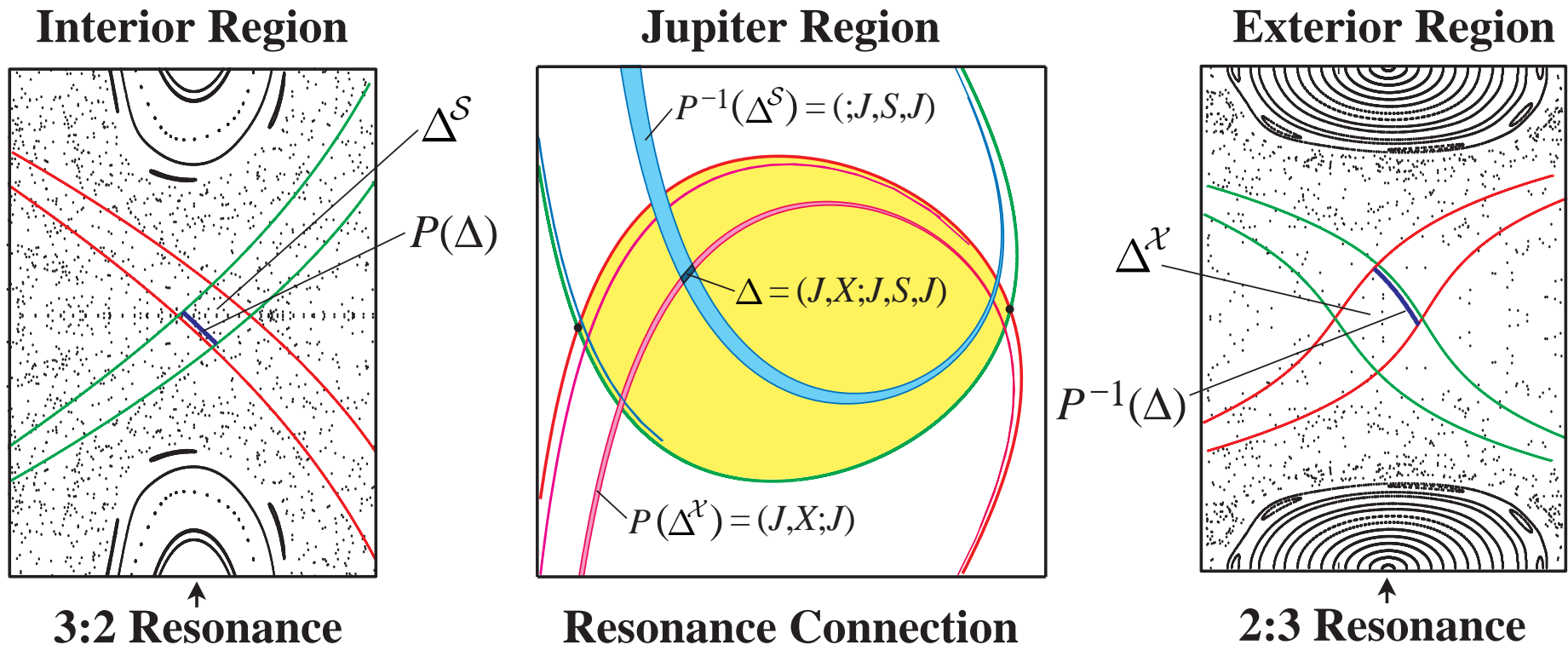
- We have referred to a heteroclinic intersection Δ connecting interior $\Delta^{\mathcal{S}}$ and exterior $\Delta^{\mathcal{X}}$ resonance intersection regions. Below, we show image of $\Delta^{\mathcal{X}}$ (2:3 resonance portion) and pre-image of $\Delta^{\mathcal{S}}$ in the J region. Their intersection $\Delta = P(\Delta^{\mathcal{X}}) \cap P^{-1}(\Delta^{\mathcal{S}})$ contains all orbits whose itineraries have central block $(J, X; J, S, J)$, corresponding to at least one transition between the exterior 2:3 resonance and interior 3:2 resonance.

Poincare Section in the Jupiter Region



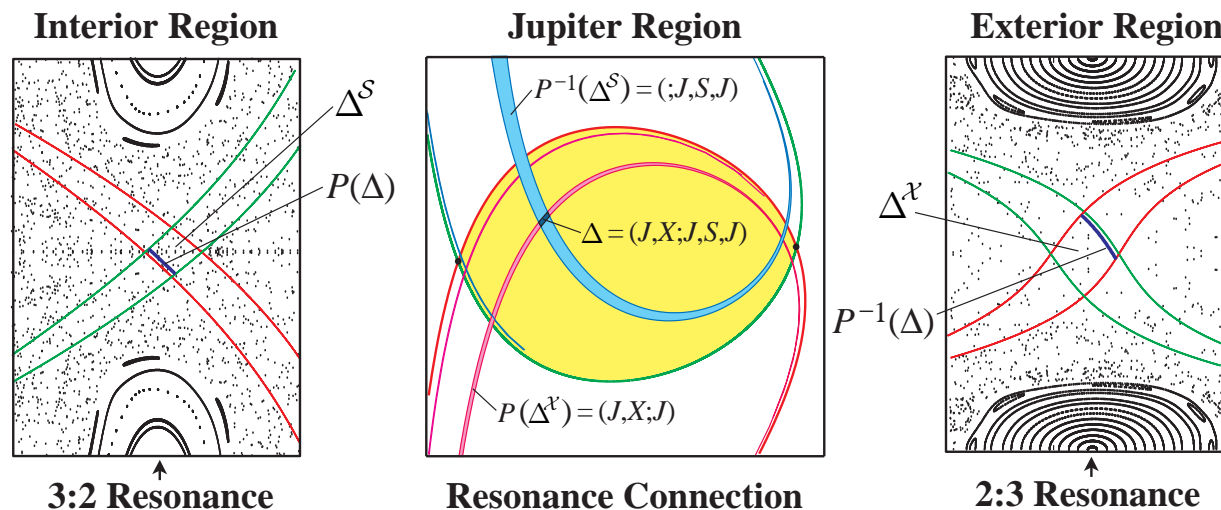
■ Connection Between Interior and Exterior Resonances

- Δ contains orbits in *transition* between the 2:3 to 3:2 resonances.
- Comets such as *Oterma* have passed through analogous regions.



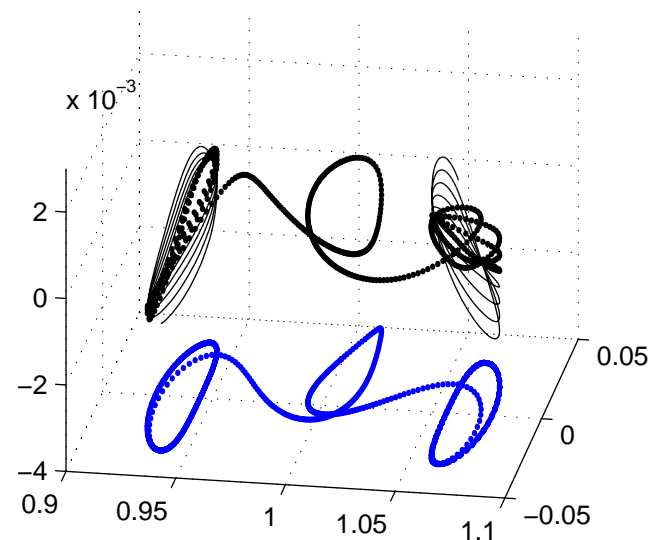
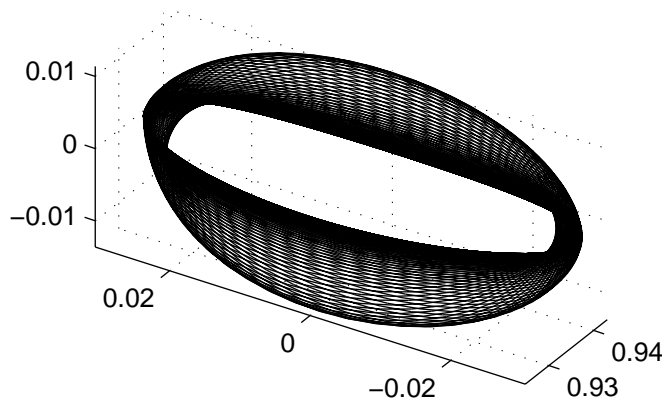
■ Resonance Connection for Three Degrees of Freedom

- It is reasonable to conclude that, within the full three-dimensional model, *Oterma's* orbit lies in an analogous region of phase space.
- It is therefore within the L_1 and L_2 periodic and quasiperiodic orbit manifold *tubes*, whose complex global dynamics lead to *intermittent behavior*, including *resonance transition*.
- More study is needed for a thorough understanding of the resonance transition phenomenon. The tools developed in this course (*dynamical channels*, *symbolic dynamics*, etc.) should lay a firm theoretical foundation for any such future studies.



■ Future Work: Extension to Three Dimensions

- ▶ Natural extension: apply same methodology to ***3D CR3BP***.
- Seek homoclinic & heteroclinic orbits associated with 3D periodic “halo” & quasi-periodic “quasi-halo” & Lissajous orbits about L_1 & L_2 . Dimension count suggests heteroclinic intersections exist.
- Union would be 3D homoclinic-heteroclinic chains around which symbolic dynamics could be used to track a variety of exotic orbits.
- ***Three-dimensional dynamical channels*** will provide more complete understanding of phase space transport mechanisms.

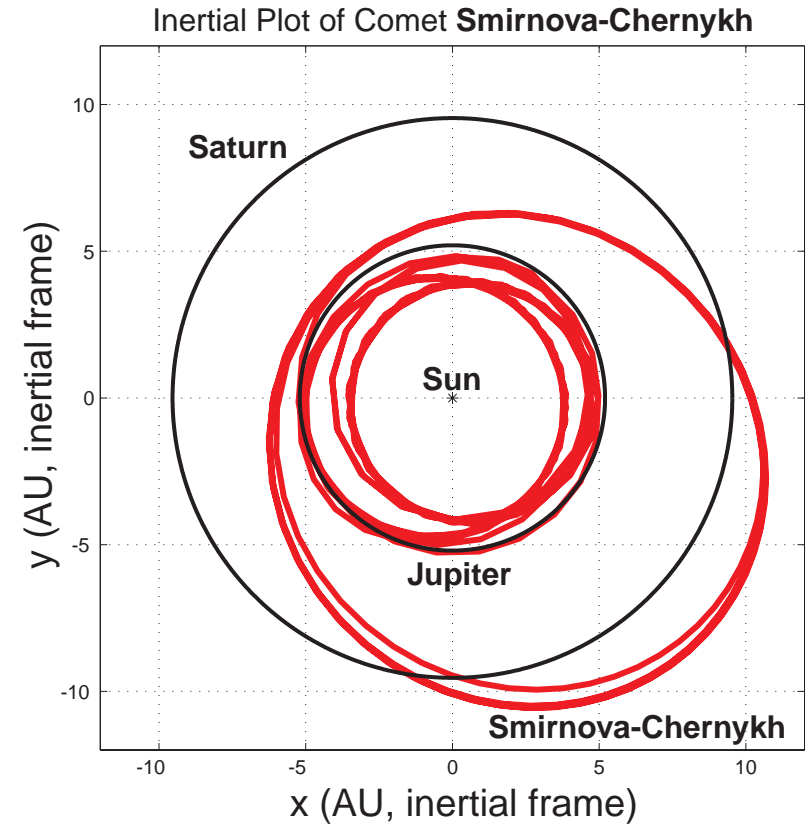
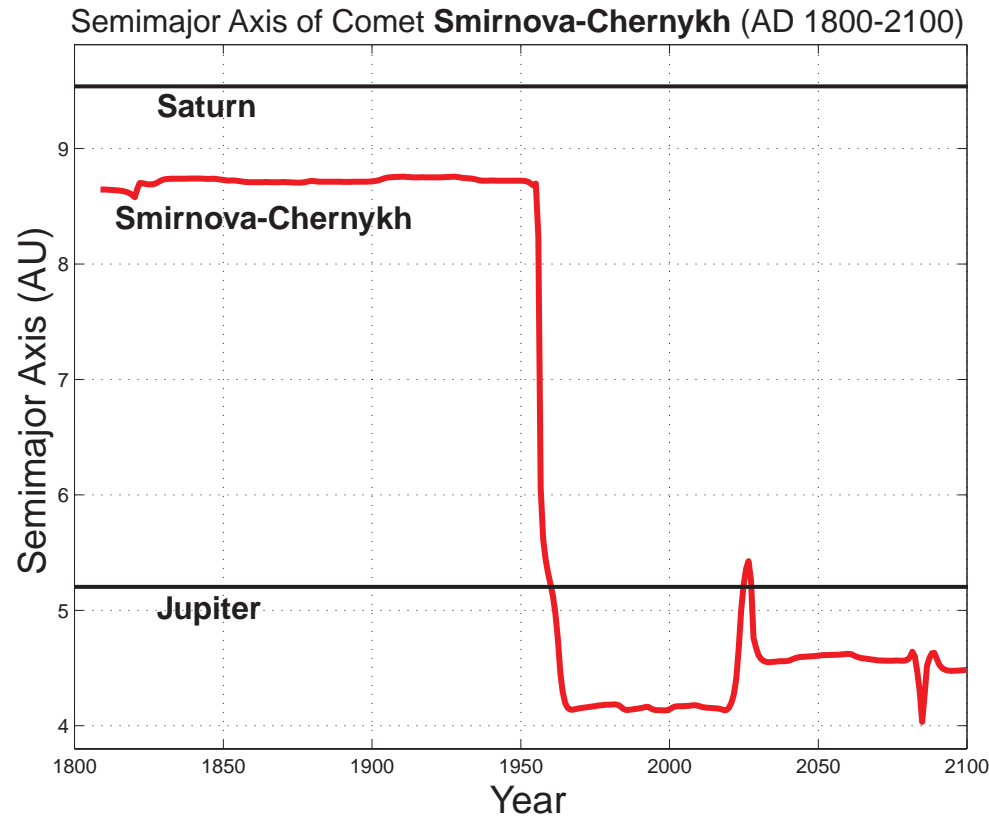


■ Future Work: Coupling of Two 3-Body Systems

- ▶ Dynamics governing *transport between adjacent planets*.
 - Coupled 3-body problem: e.g., comet between Jupiter & Saturn.
 - Between the two planets, the comet's motion is mostly heliocentric, but is precariously poised between two competing three-body dynamics.
 - In this region, heteroclinic orbits connecting Lyapunov orbits of the two different three-body systems may exist, leading to complicated transfer dynamics between the two adjacent planets.

Comet Transition Between Jupiter and Saturn

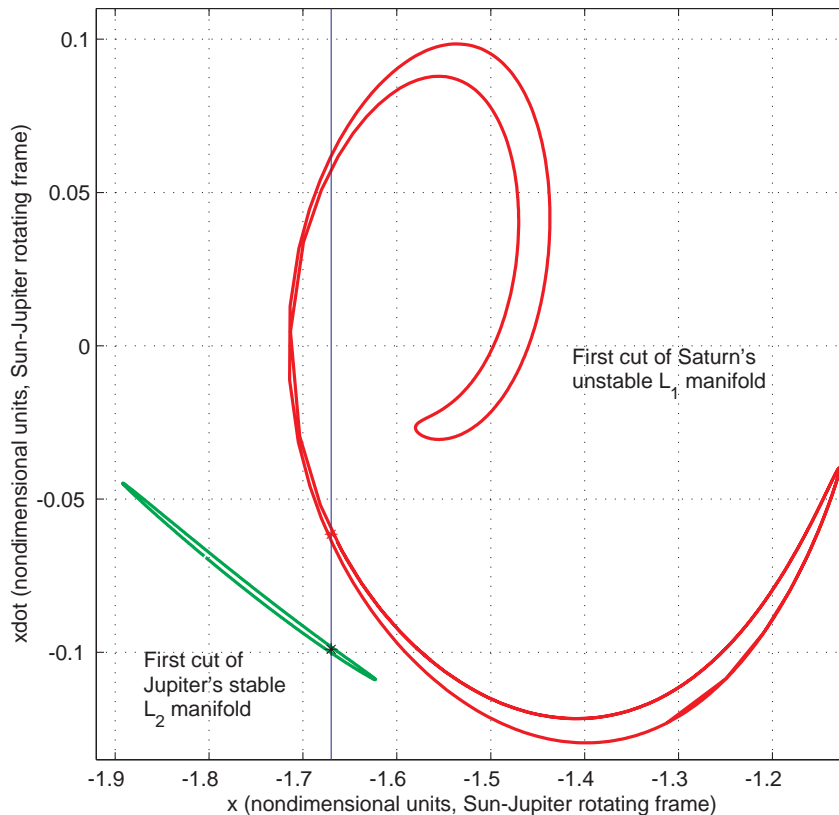
- ▶ Example: Comet *Smirnova-Chernykh* undergoes a rapid transition from **Saturn's** control to **Jupiter's** control.



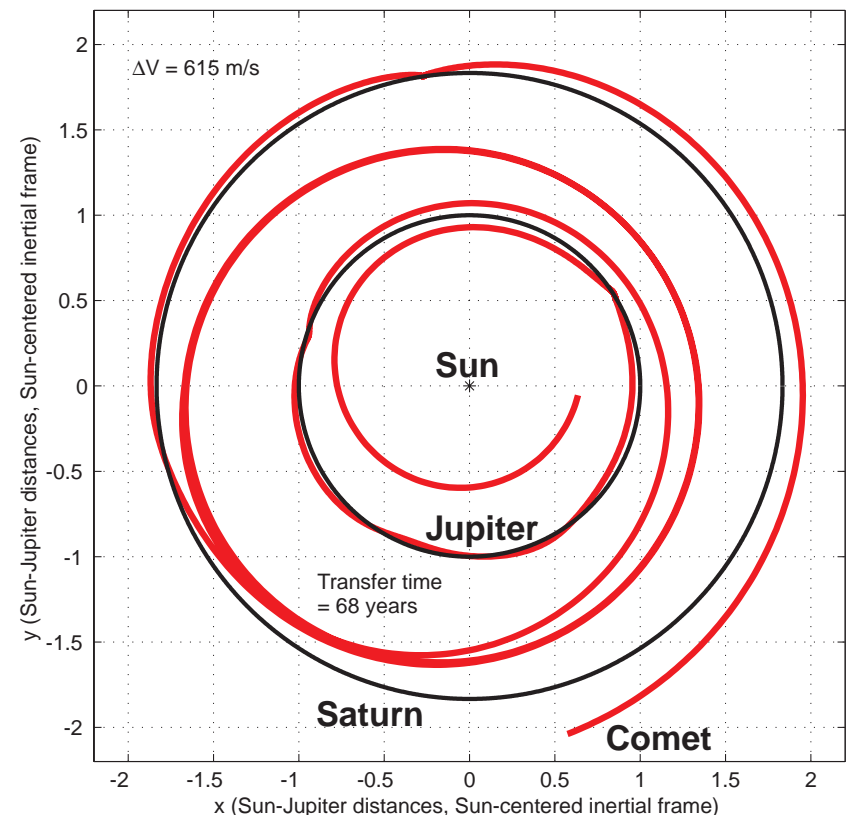
Comet Transition Between Jupiter and Saturn

- ▶ Coupled PCR3BP shows near *intersections* between Lyapunov orbit manifold tubes of Jupiter and Saturn (requiring mild ΔV).
- Natural continuous thrust of comet *outgassing* may be enough.
- Longer time integration will likely reveal genuine intersections.

Poincare section ($y=0, x<0$) in Sun-Jupiter rotating frame showing close approach of Jupiter and Saturn manifolds

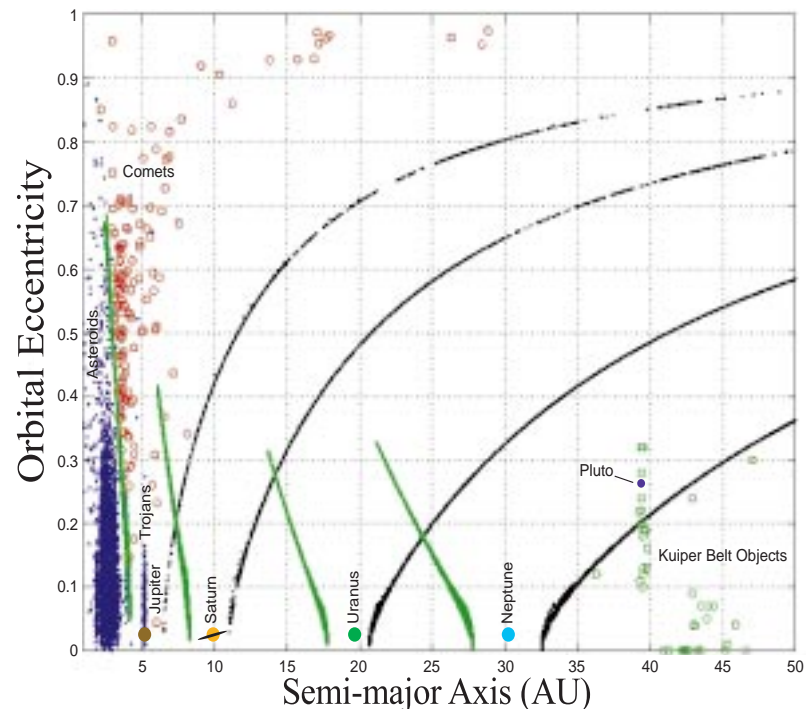


Transfer from Saturn to Jupiter via equil. region manifolds highway



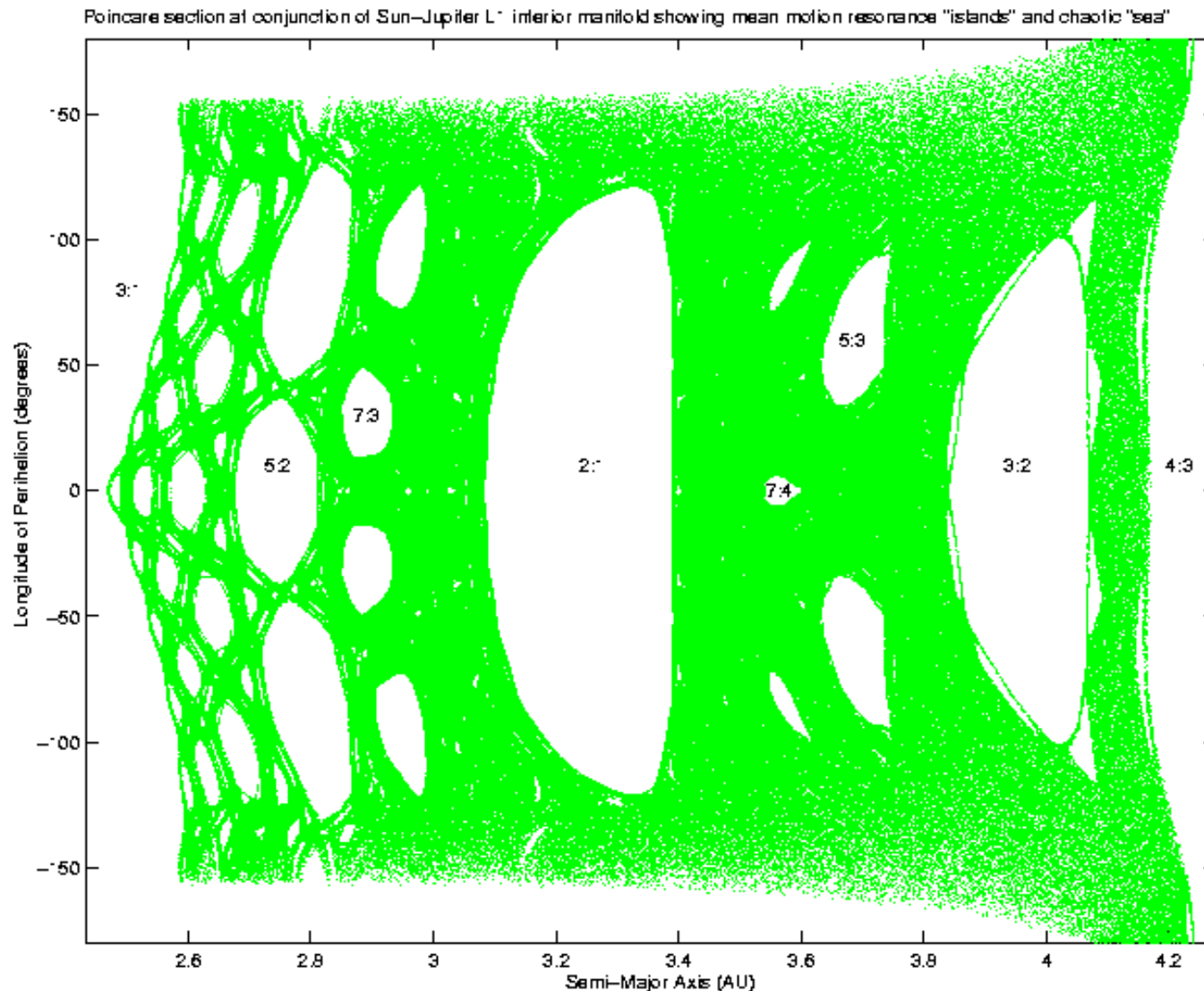
■ Future Work: Long Time Integration

- ▶ Results limited thus far to short time (a few periods of Jupiter).
- **Long time** integration (millions of Jupiter periods) will reveal *statistical* information and *new phenomena*.
- Preliminary results suggest manifold structures associated to L_1 and L_2 have helped sculpt the *solar system* and *transport material* between the planets.



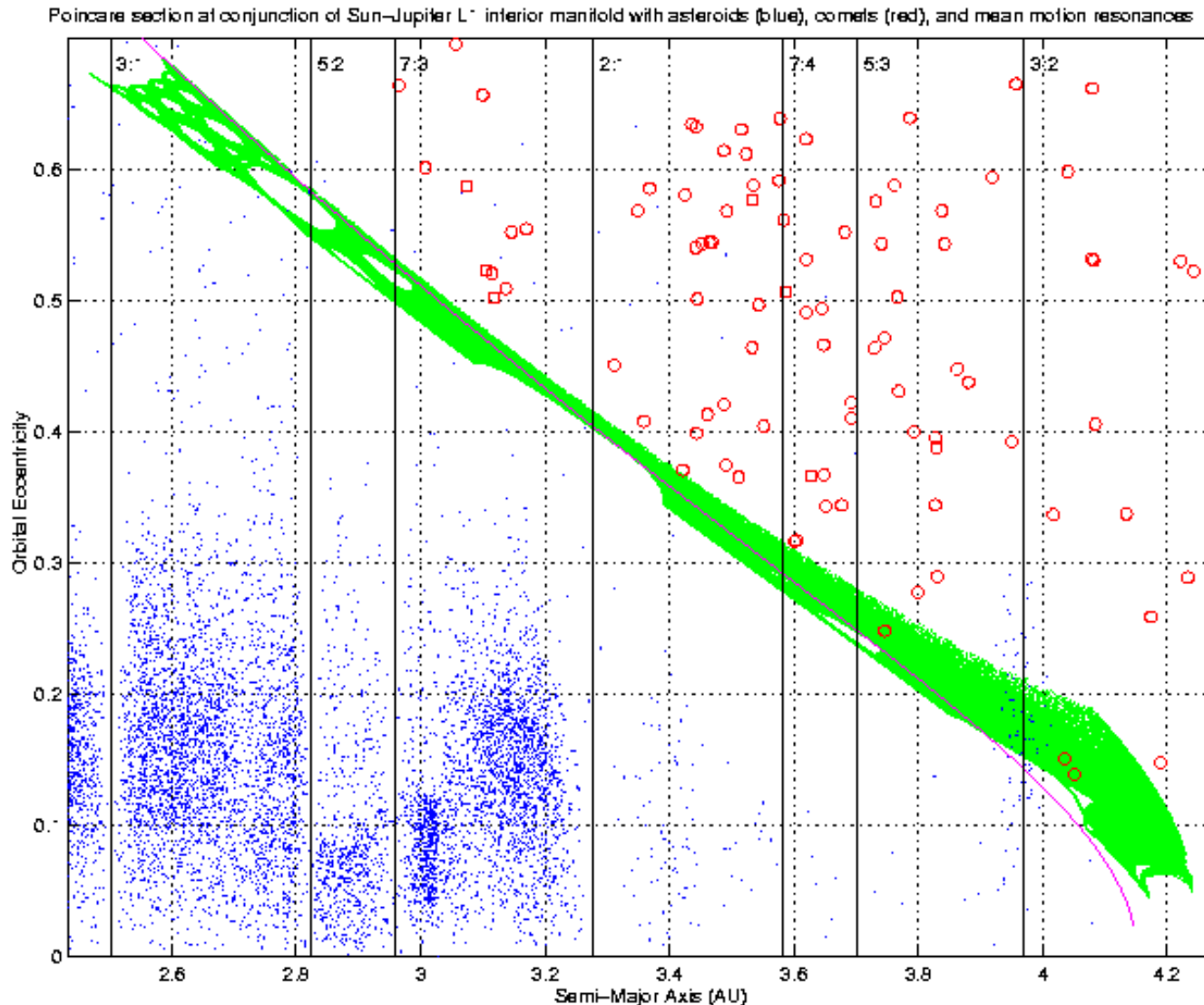
■ Long Time Integration: Jupiter's L_1 Manifolds

- We show U_1 Poincaré section of Jupiter's L_1 stable & unstable manifolds for one million iterations (in a vs. \bar{g}).



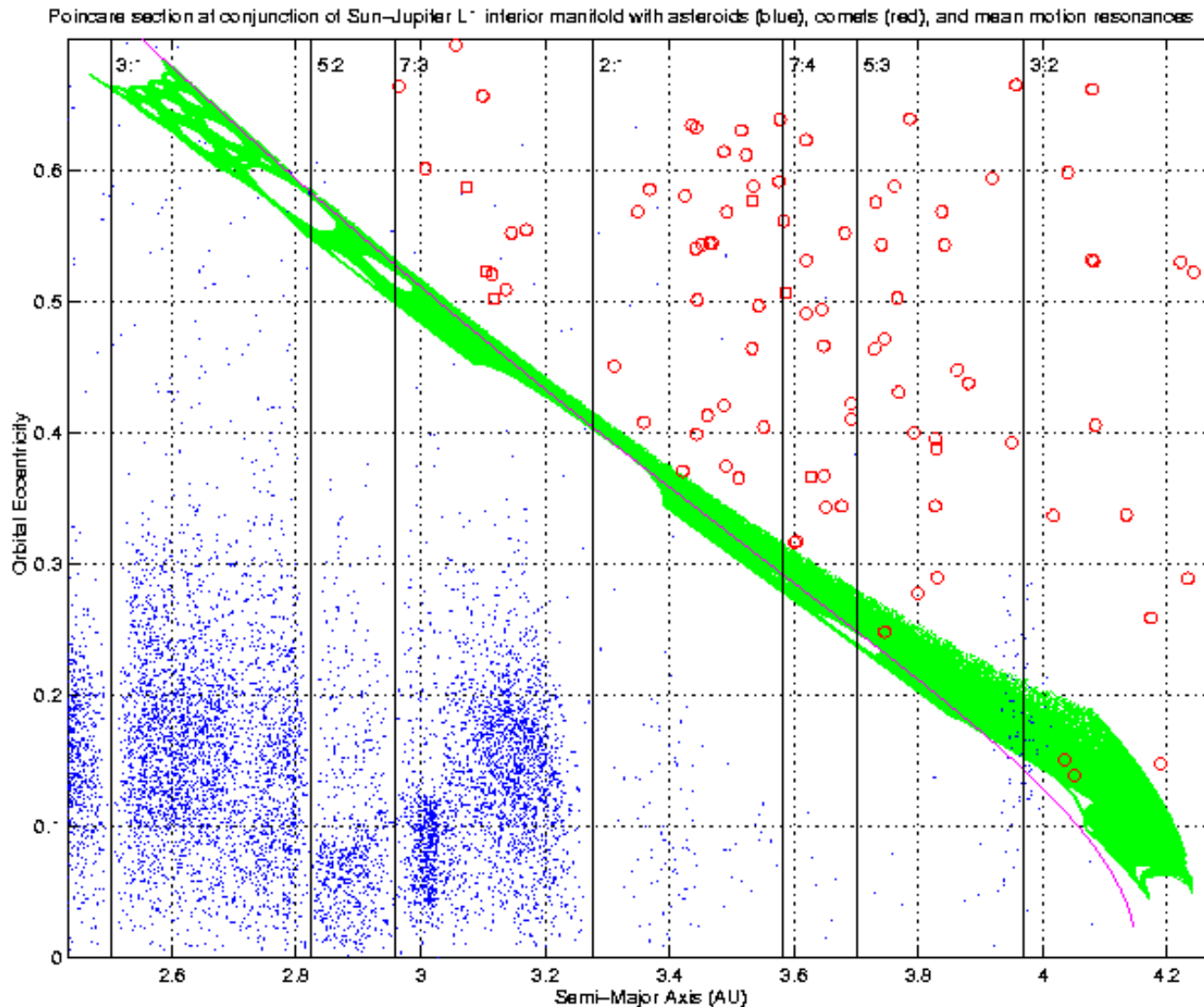
Long Time Integration: Jupiter's L_1 Manifolds

- We can also plot this in semimajor axis a vs. eccentricity e .
Away from L_1 , manifold hugs curve given by $C = \frac{1}{a} + 2\sqrt{a(1 - e^2)}$.



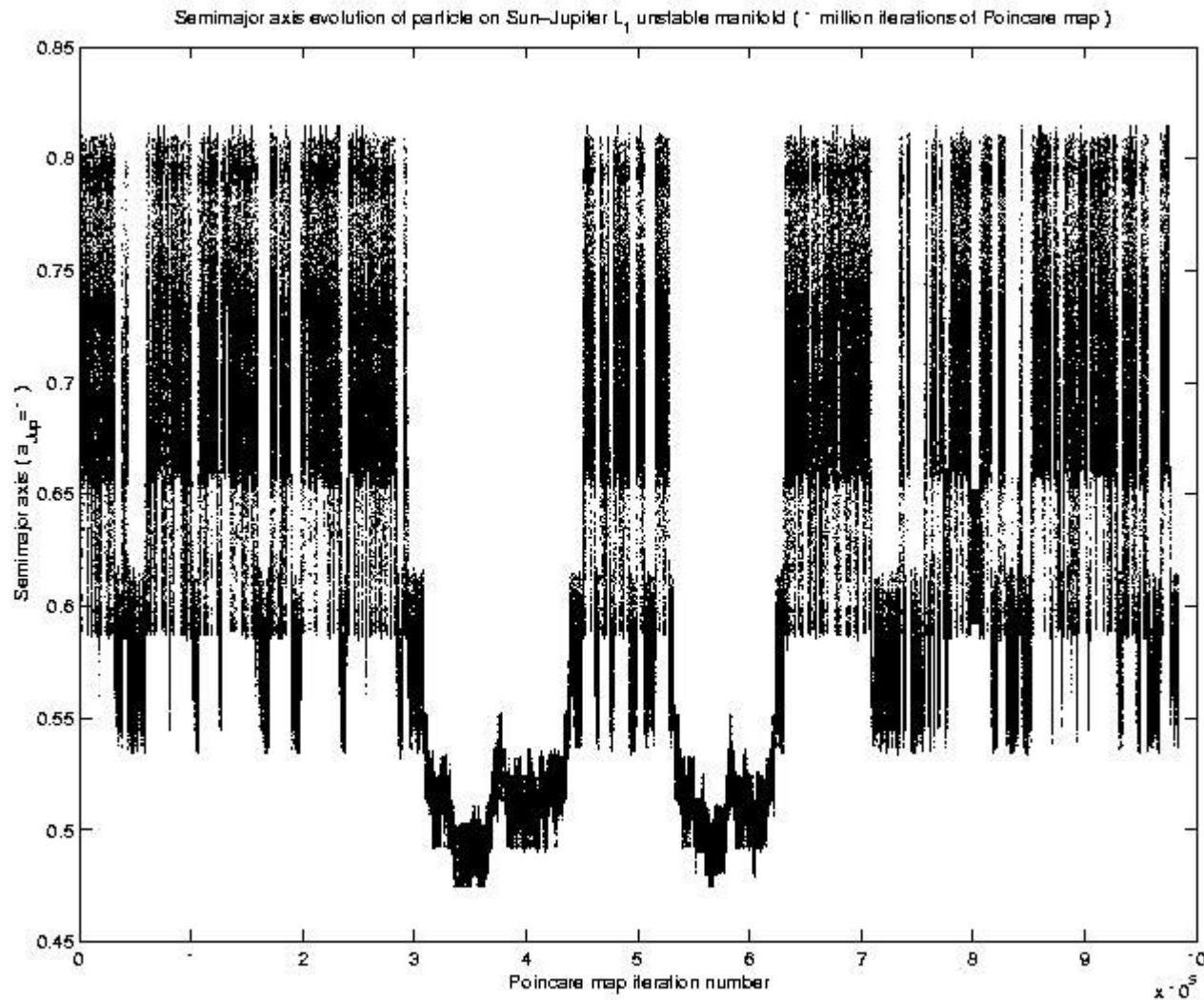
Long Time Integration: Jupiter's L_1 Manifolds

- Note how *manifold* acts as stability boundary, separating stable *asteroids* from unstable *comets*.



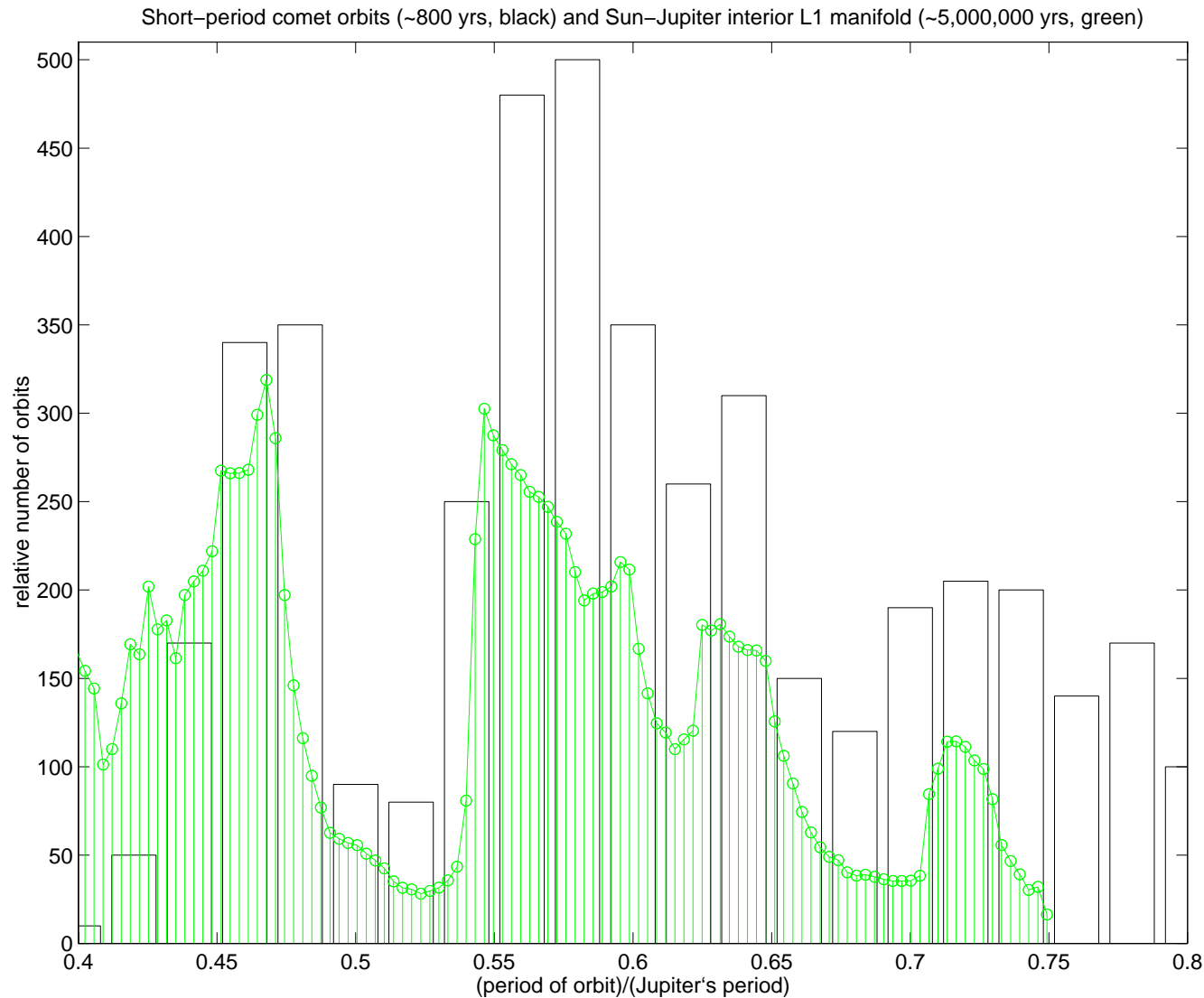
■ Intermittent Behavior Along Jupiter's L_1 Manifolds

- Time history of semimajor axis a for one million iterations shown.
- Manifold exhibits *intermittency*, *jumping*, *sticking*.



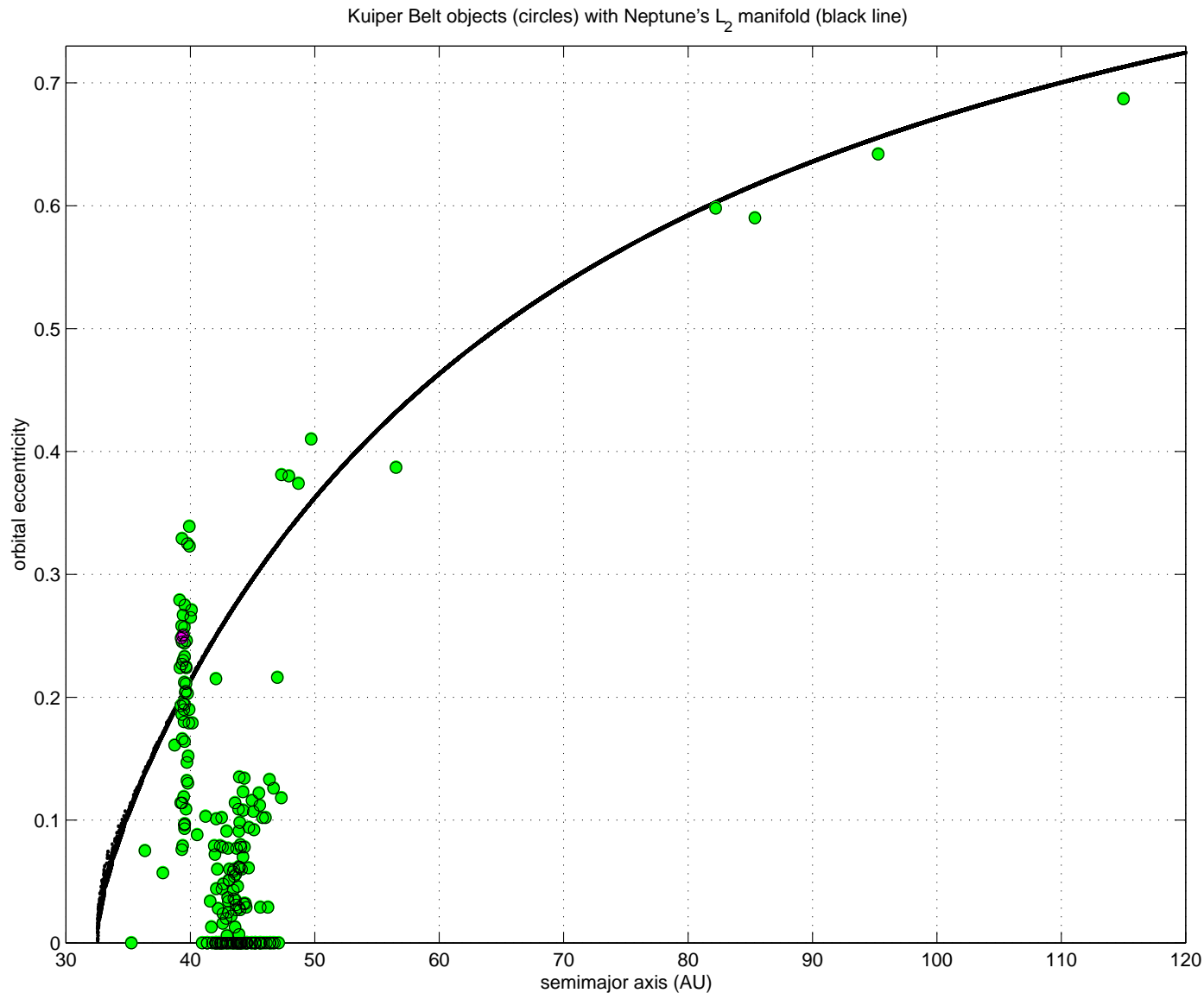
Comet Distribution Matches Jupiter's L_1 Manifolds

- Taking histogram of Jupiter's L_1 manifolds shows fair agreement with distribution of Jupiter comets. Same dynamics is at work.



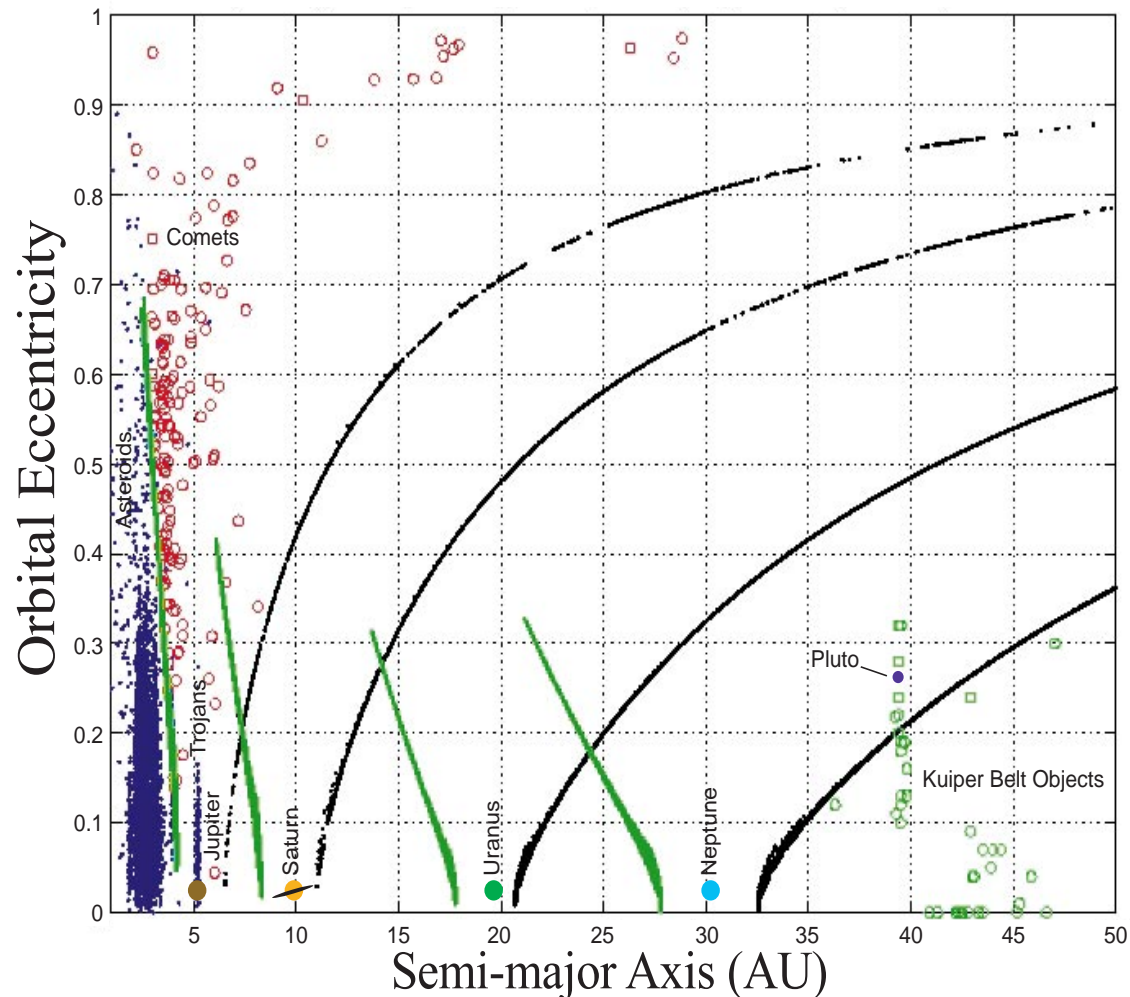
Kuiper Belt and Neptune's L_2 Manifolds

- Just as Jupiter's manifolds determine asteroid & comet distribution and transport, Neptune's manifolds may govern the Kuiper belt.



Transport Between Asteroid Belt and Kuiper Belt

- Intersections between L_1 and L_2 manifold structures between adjacent planets may provide a “highway” connecting the asteroid and Kuiper belts, where material can collect.



■ Conclusion: L_1 and L_2 Manifolds are Important!

- The invariant manifold structures associated to L_1 and L_2 , as well as the homoclinic-heteroclinic *dynamical channels* connecting them, are *fundamental tools* that can aid in understanding mechanisms of transport throughout the solar system.

